

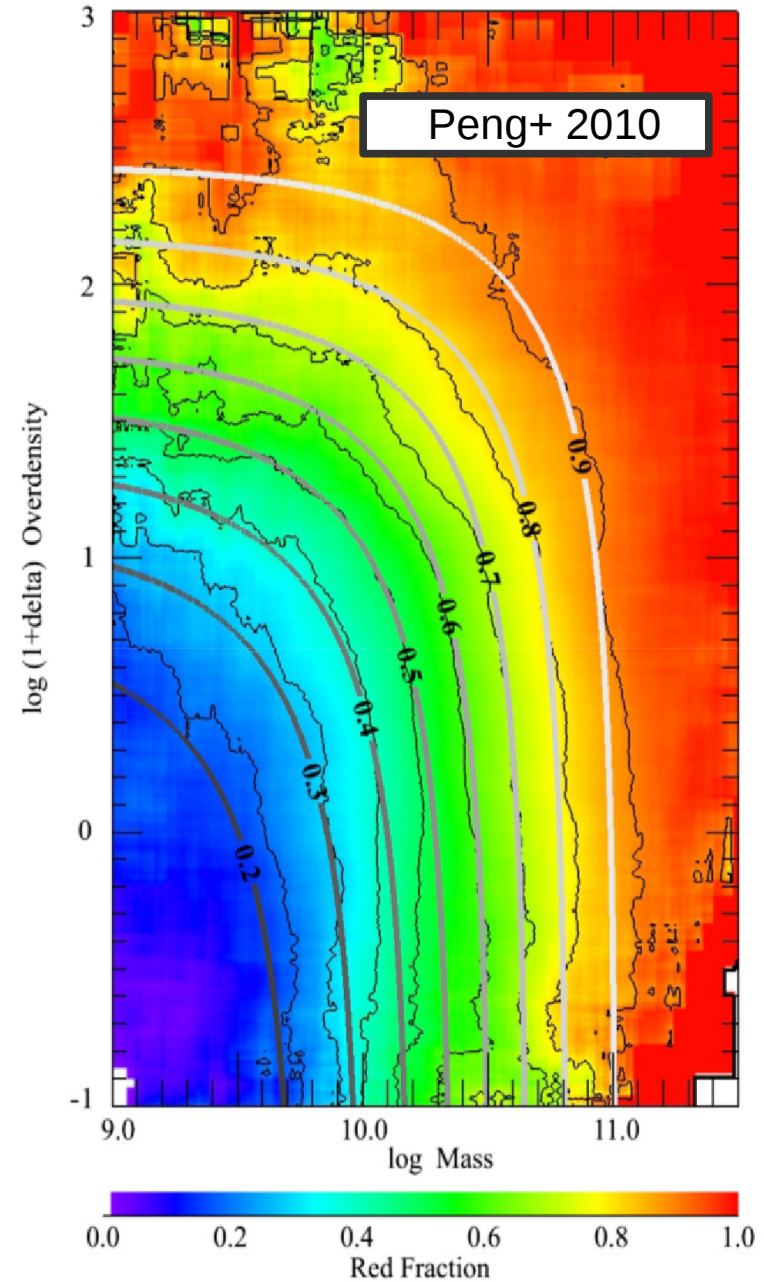
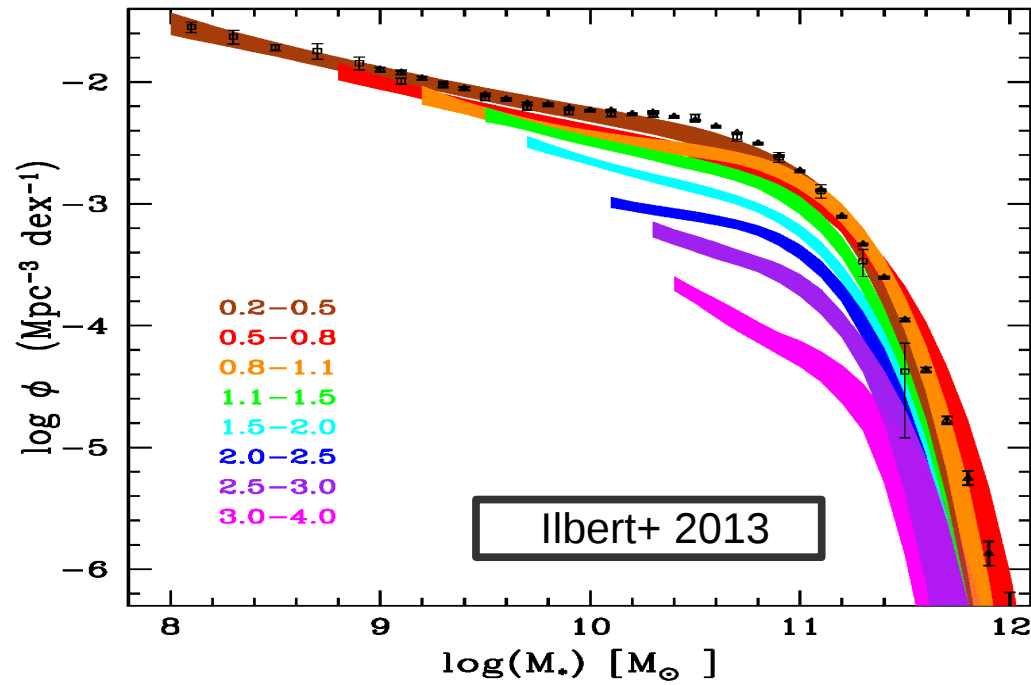
# AGN evolution from galaxy evolution viewpoint

Neven Caplar, Simon J. Lilly, Benny Trakhtenbrot

ETH Zurich

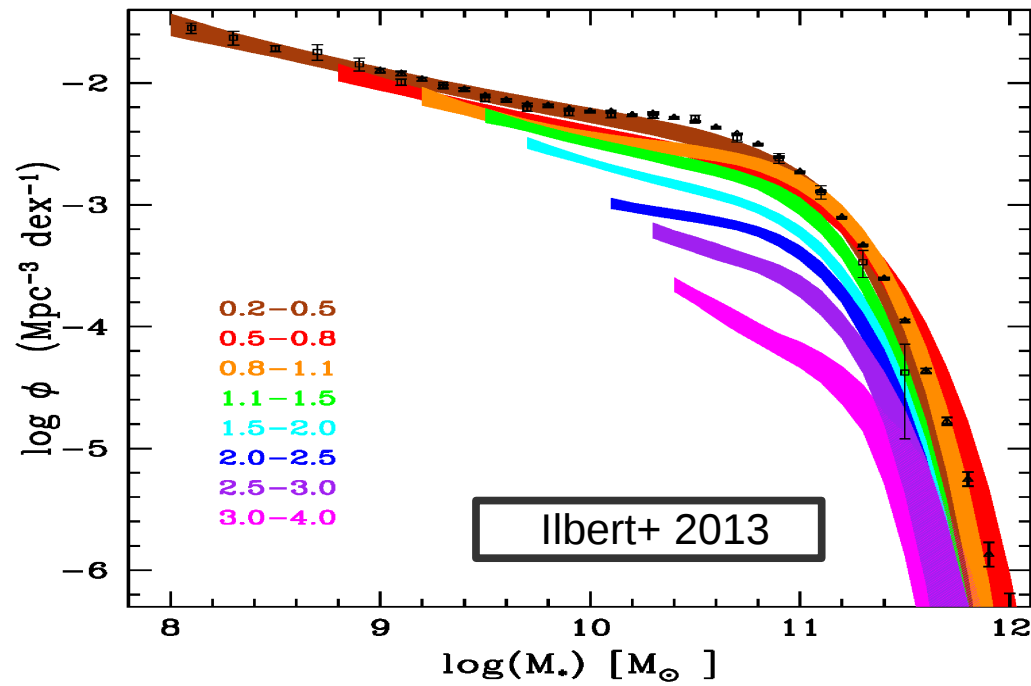
# Motivation

- Great improvements in our knowledge of galaxy population

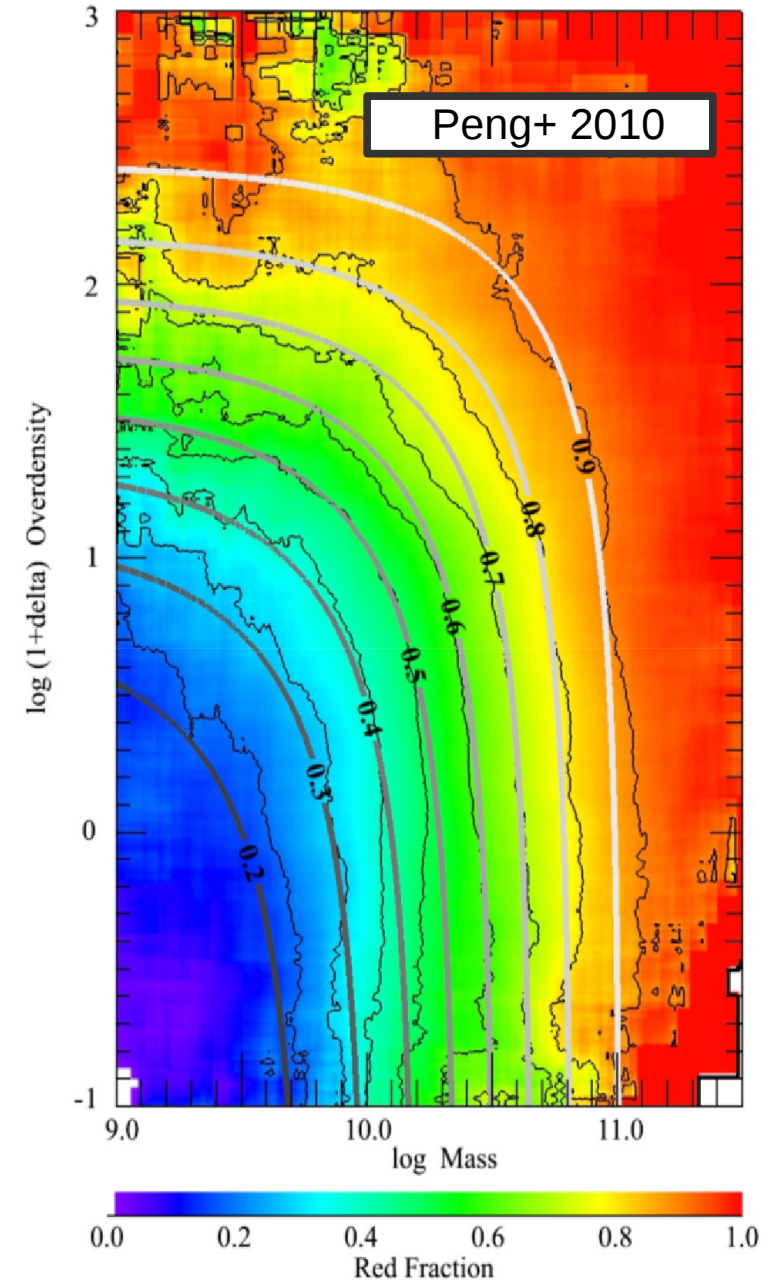


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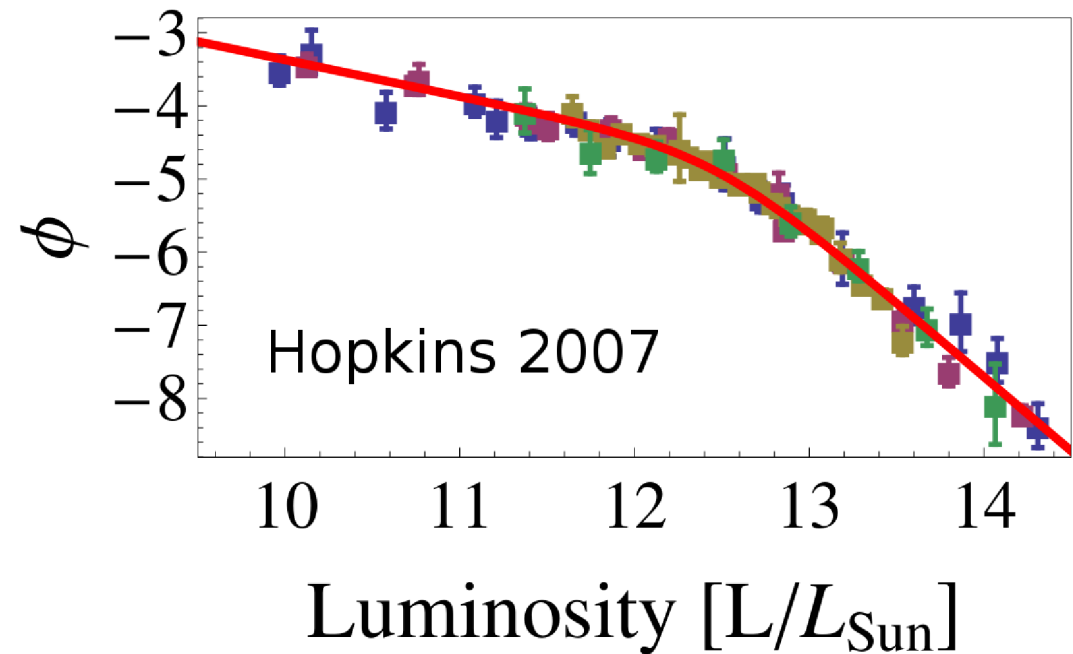
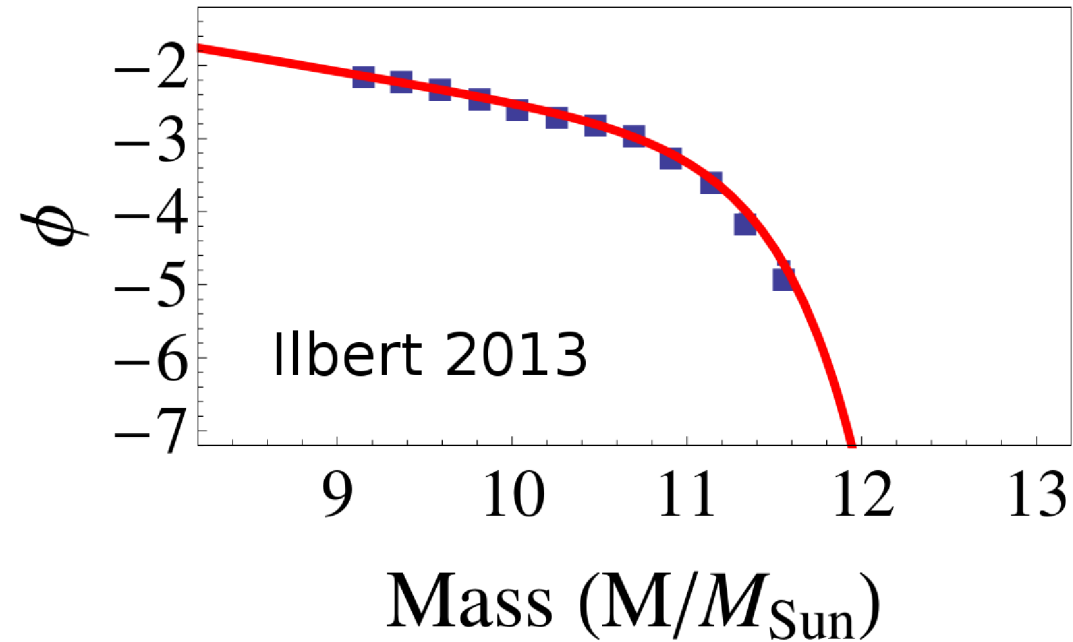


- We wish to describe AGN population with a model which is
  - Phenomenological
  - Analytical
  - Simple
  - Data driven



# Outline

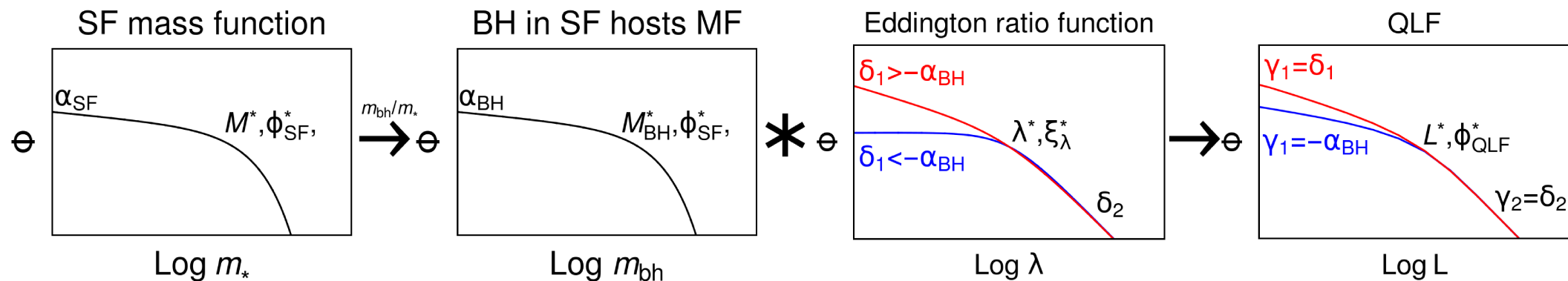
- What we can learn just from evolution of quasar luminosity function
  - How do get connect quasar luminosity function and galaxy mass function
  - Redshift evolution of these functions
  - Connecting these evolutions
- Mass ratio ( $m_{\text{bh}}/m_*$ ) evolution
  - Hints for mass evolution
  - Observational consequences



# Quasar luminosity function is convolution of galaxy mass function and Eddington ratio function

## Ansätze

- Radiatively efficient AGNs are in star forming galaxies
- Distribution of Eddington ratio does not depend on the mass of the black hole
- Mass of central black hole proportional to stellar mass
- To make quasar luminosity function convolve
  - AGN mass function & Eddington ratio function

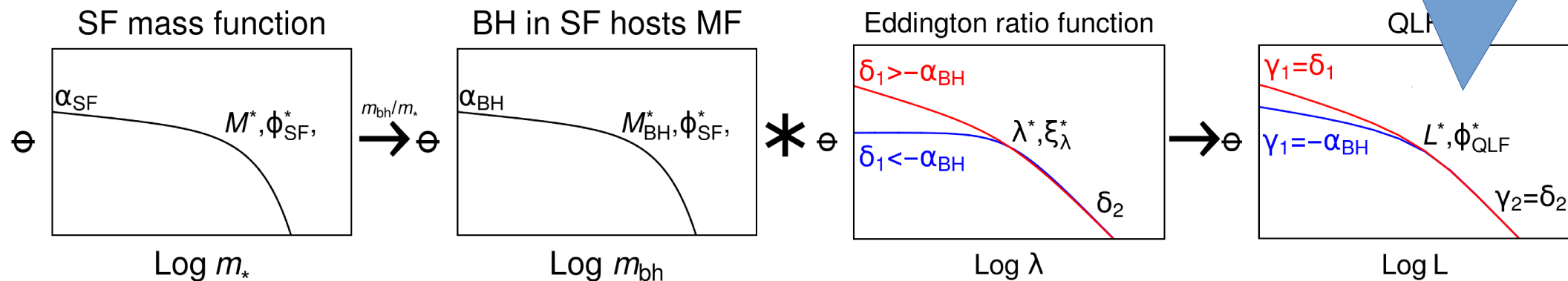


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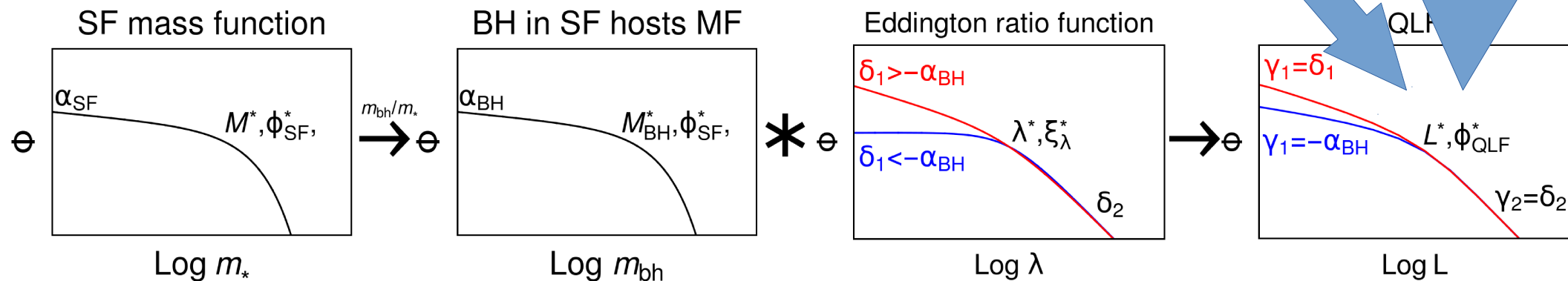
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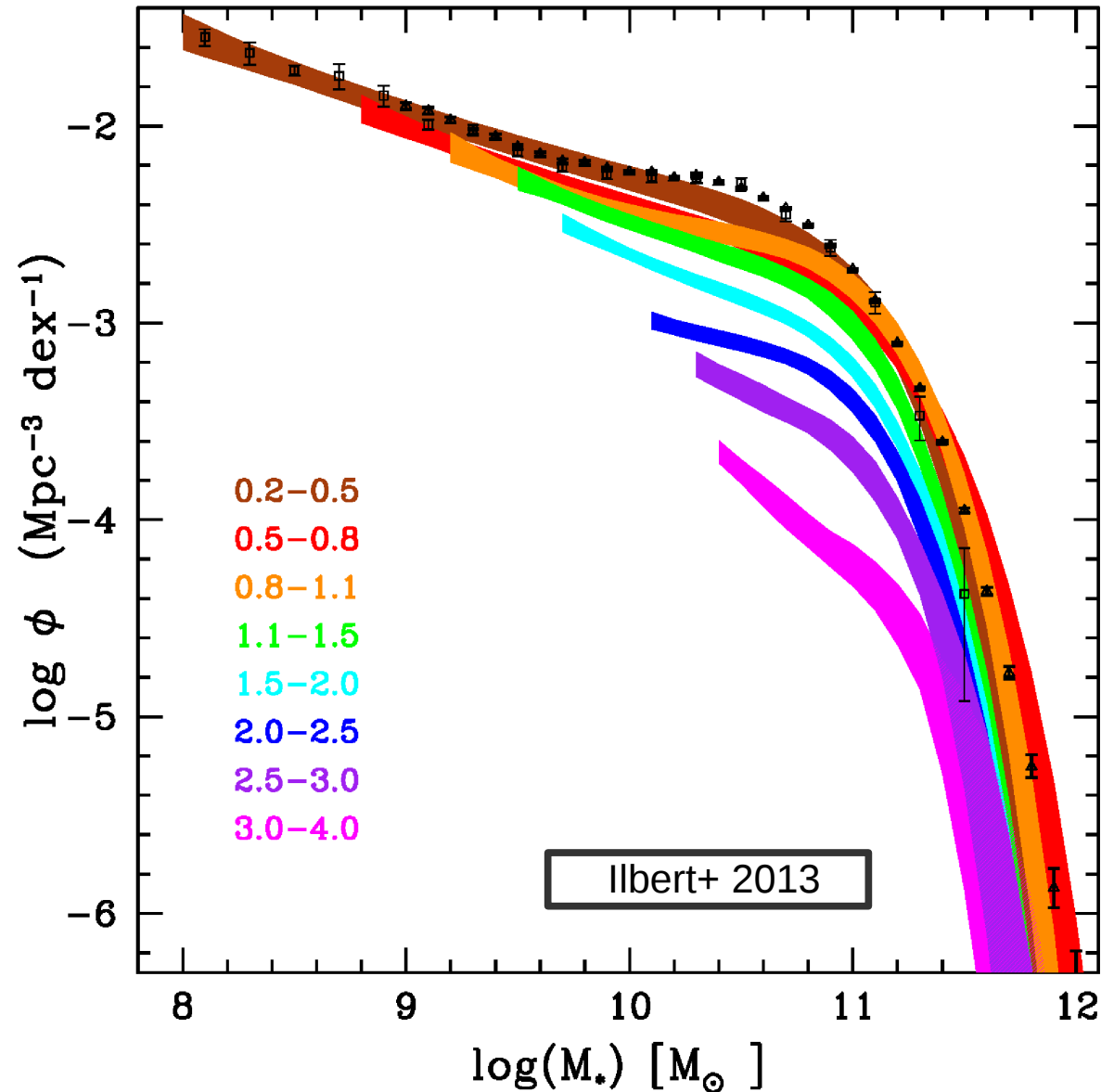
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# $M^*$ of galaxies does not change

- Fit star-forming component with

$$\phi_{SF}^* \left( \frac{m}{M^*} \right)^\alpha \exp \left[ \frac{-m}{M^*} \right]$$

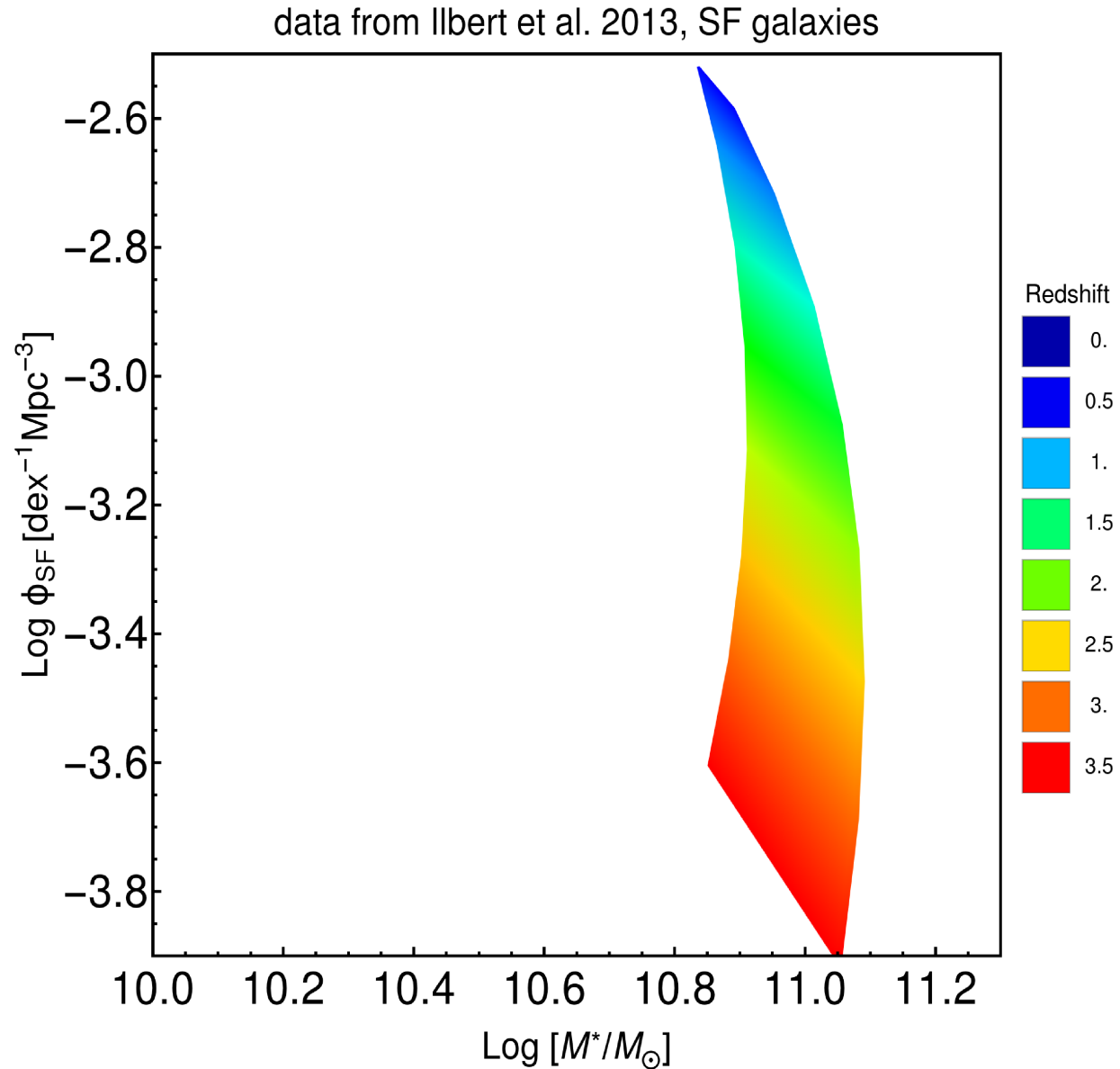




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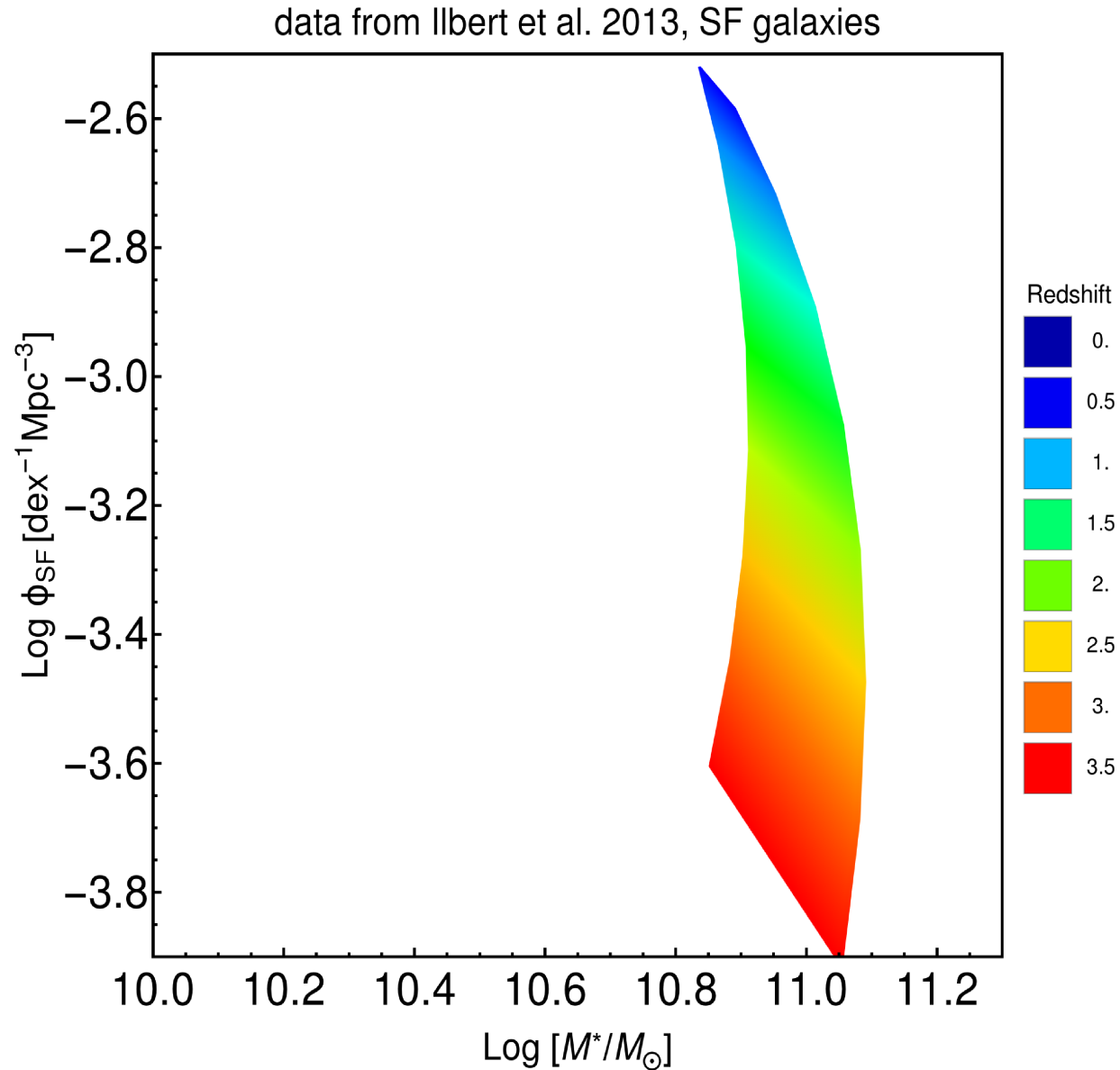


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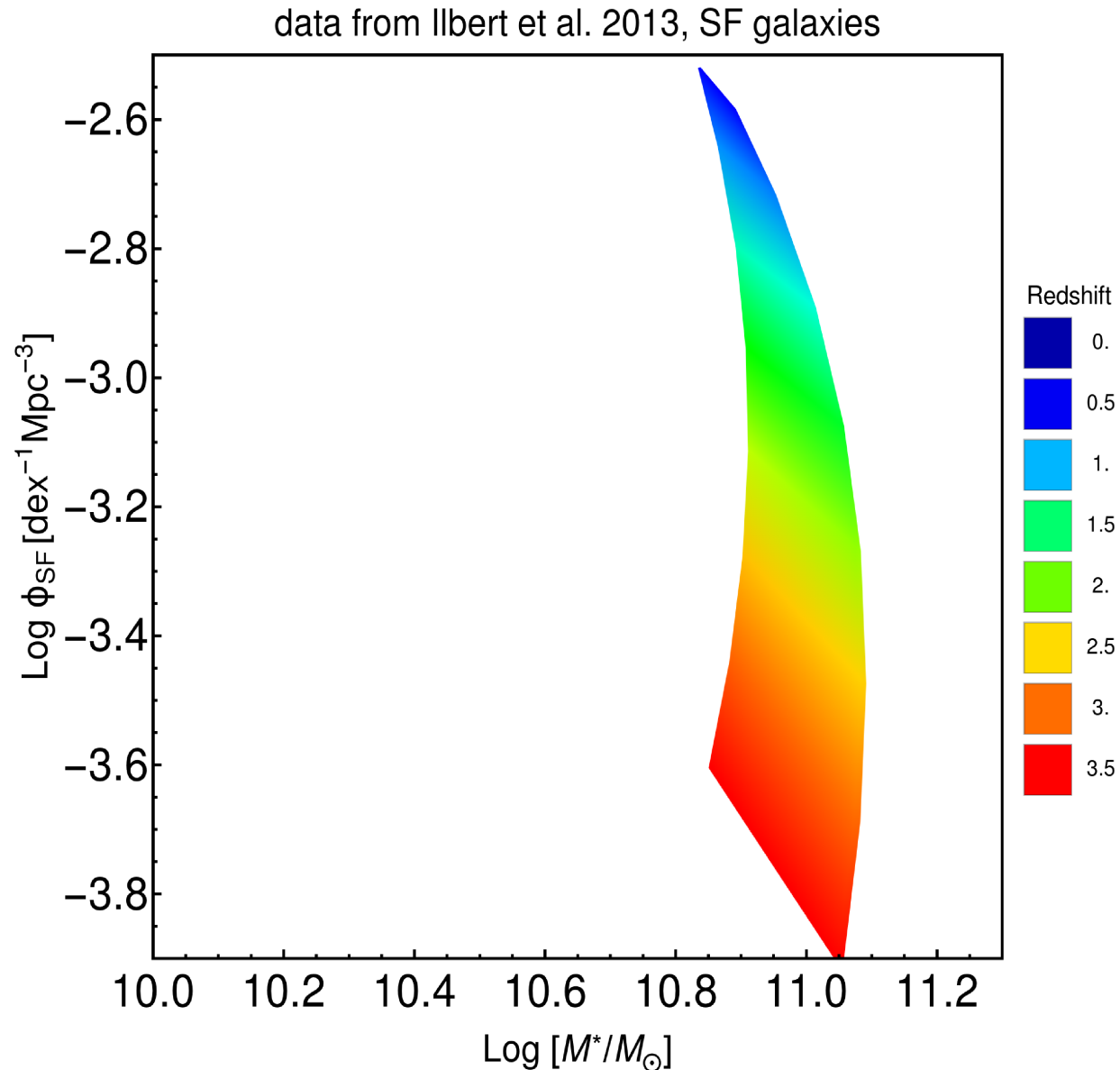


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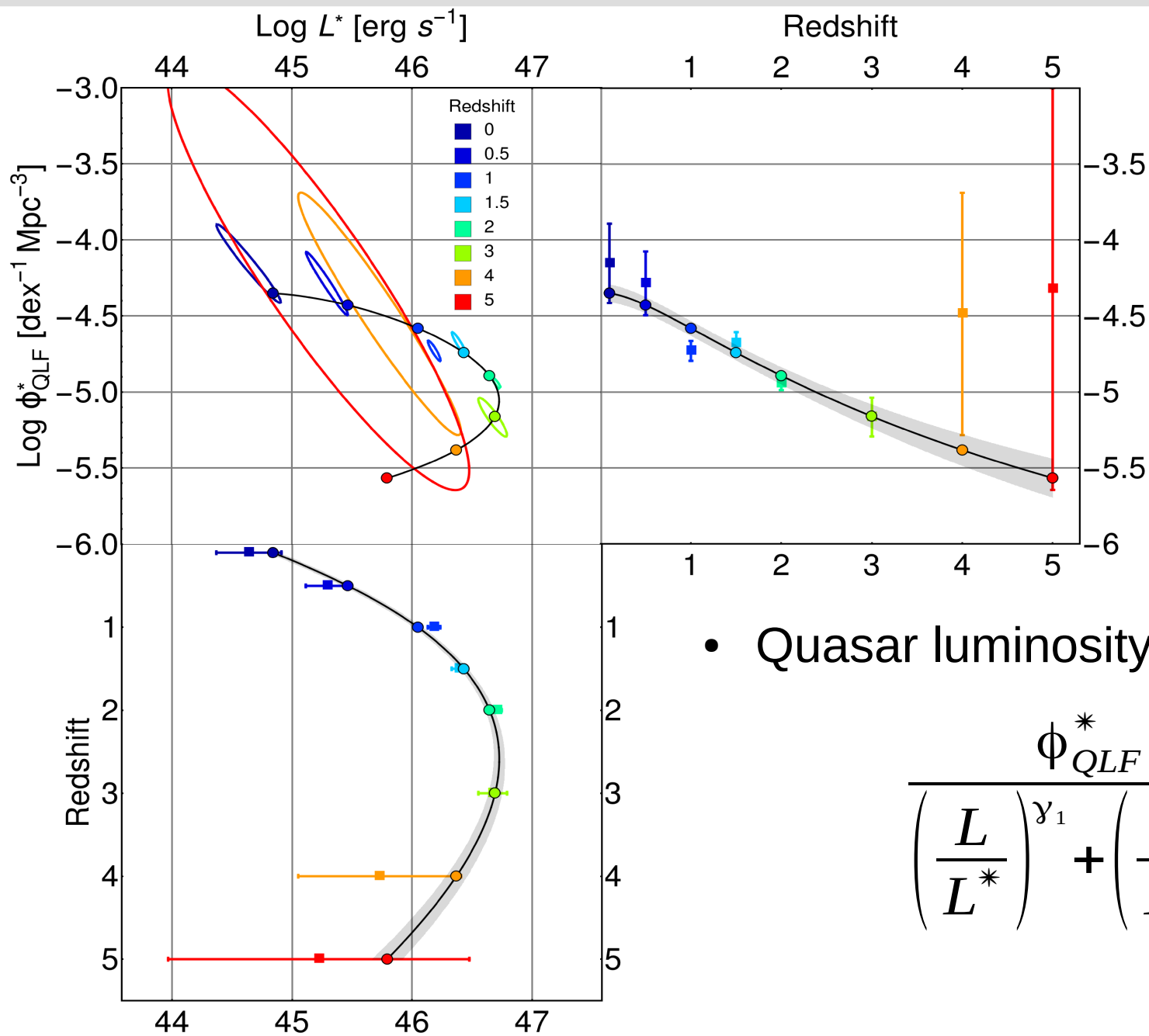
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- Minimal change of M\* up to until at least z = 3
- Normalization change consistent with simple phenomenological model for galaxies (Peng+ 2010)



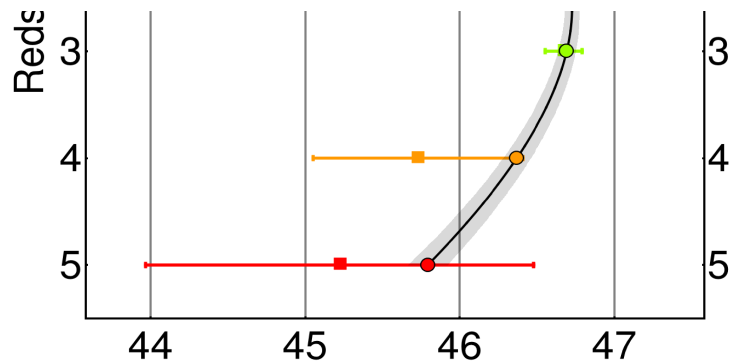
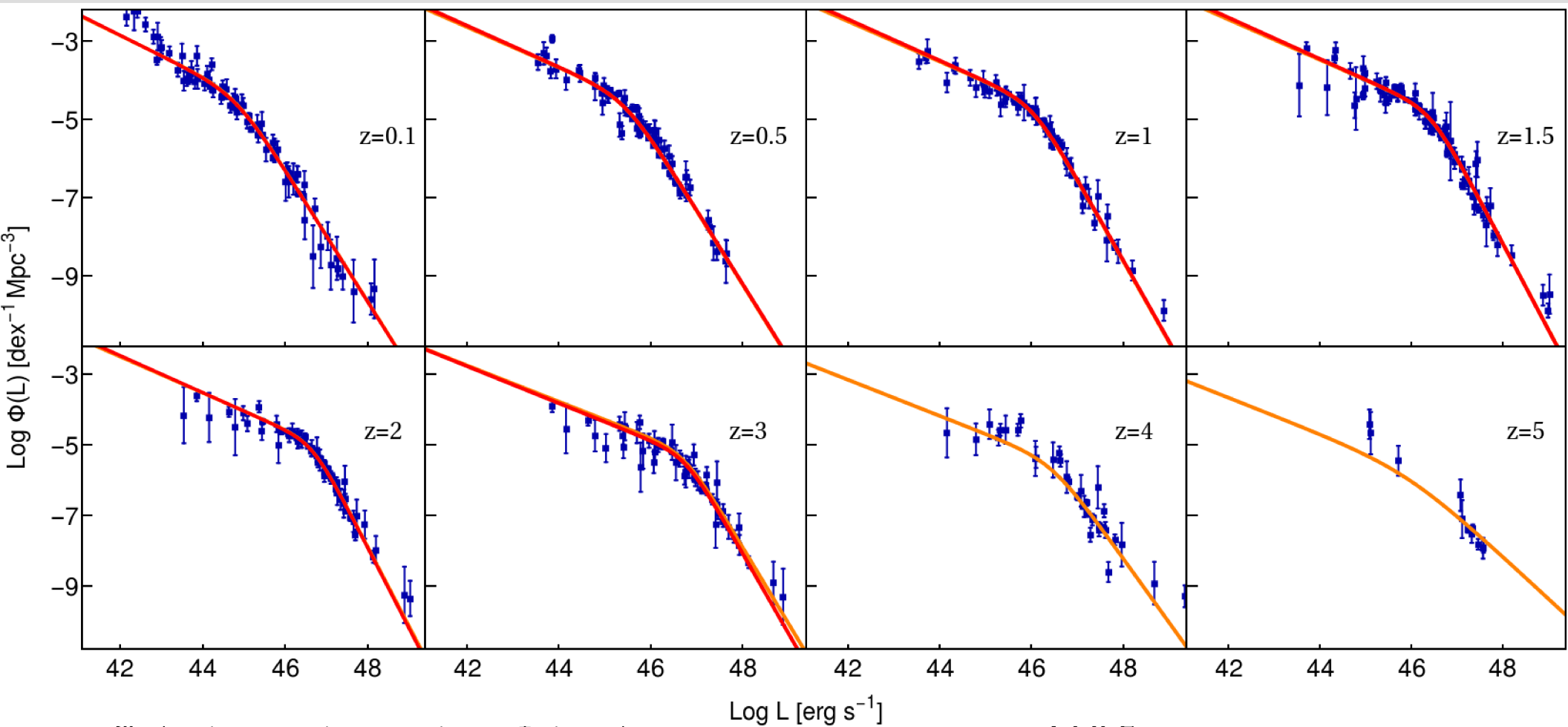
# Two interesting results from quasar luminosity function



- Quasar luminosity function

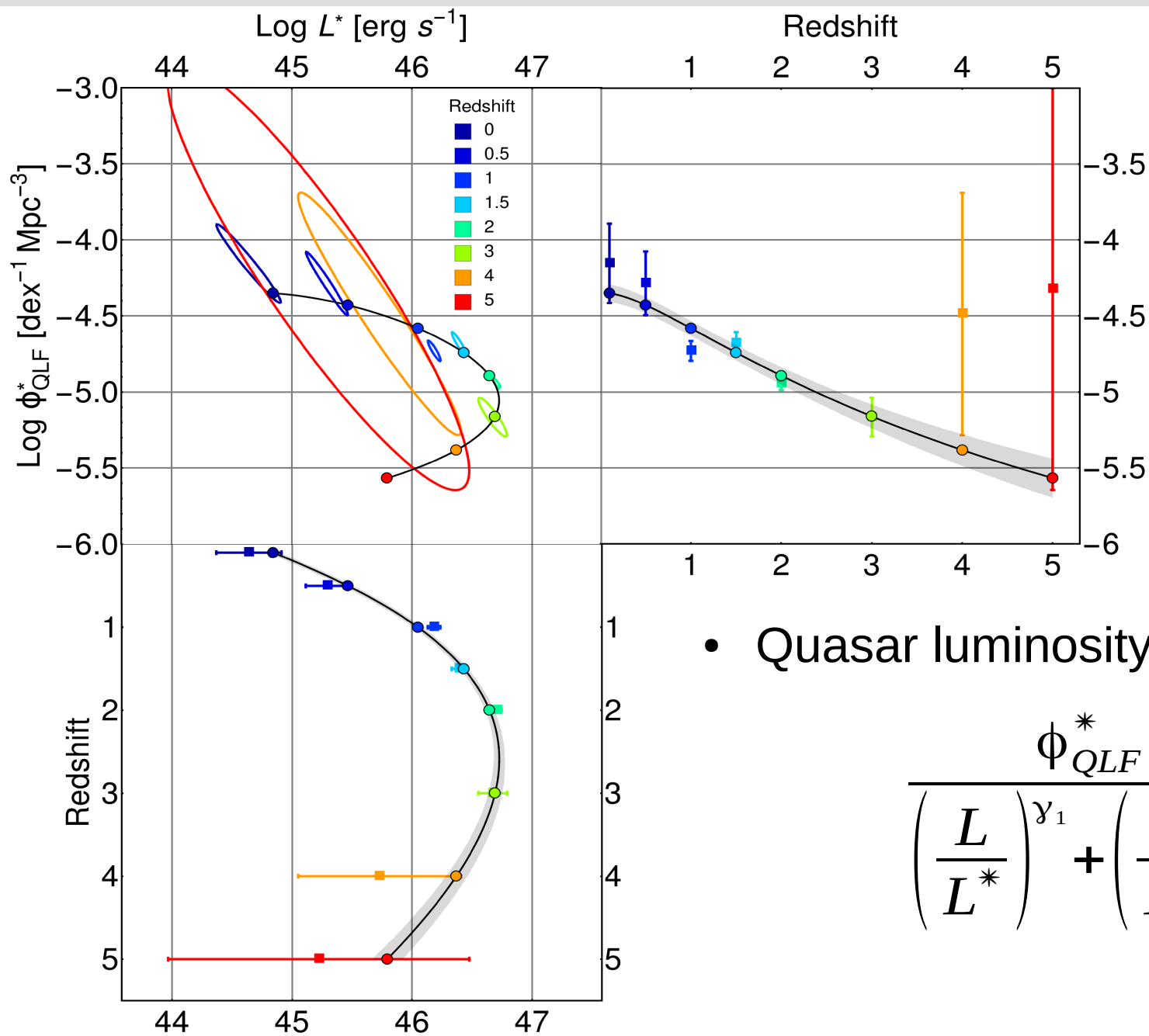
$$\frac{\phi_{QLF}^*}{\left(\frac{L}{L^*}\right)^{\gamma_1} + \left(\frac{L}{L^*}\right)^{\gamma_2}}$$

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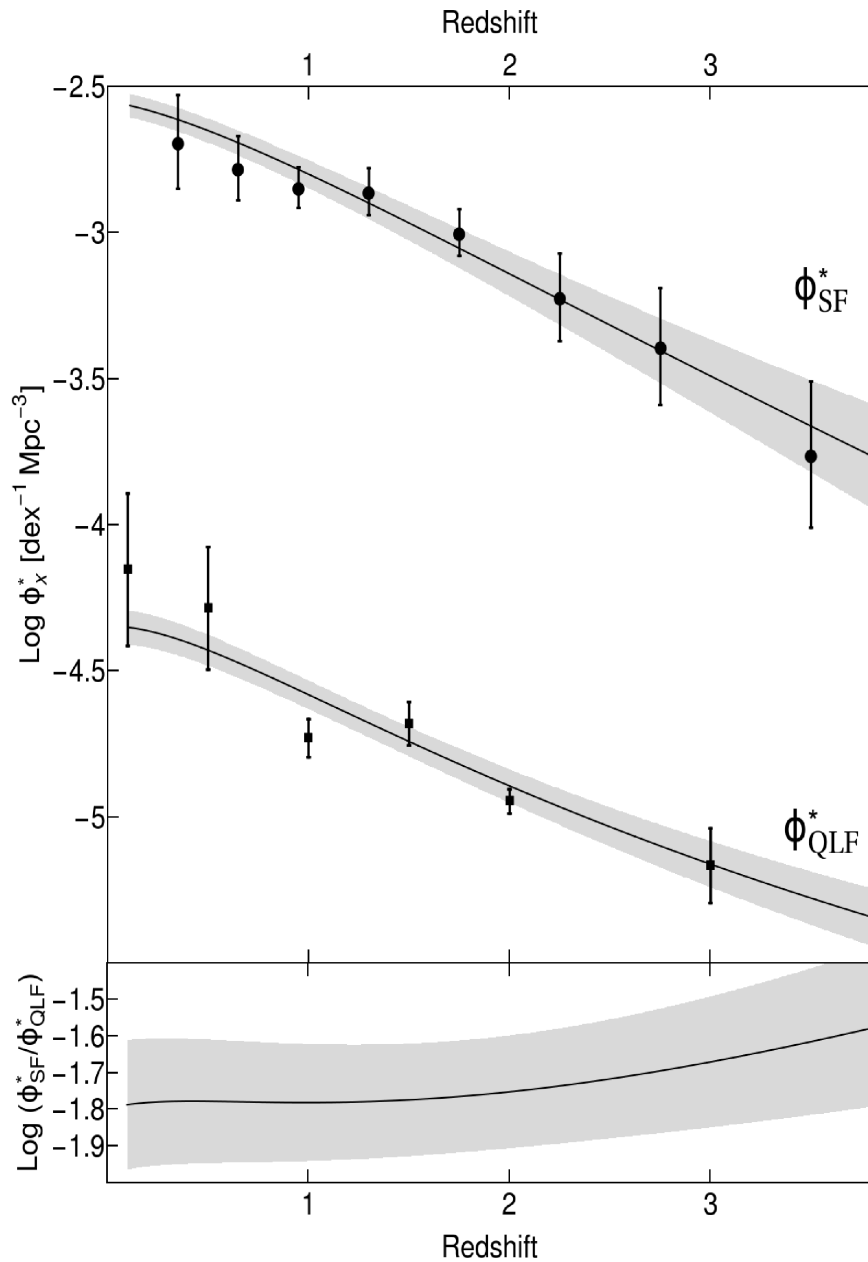
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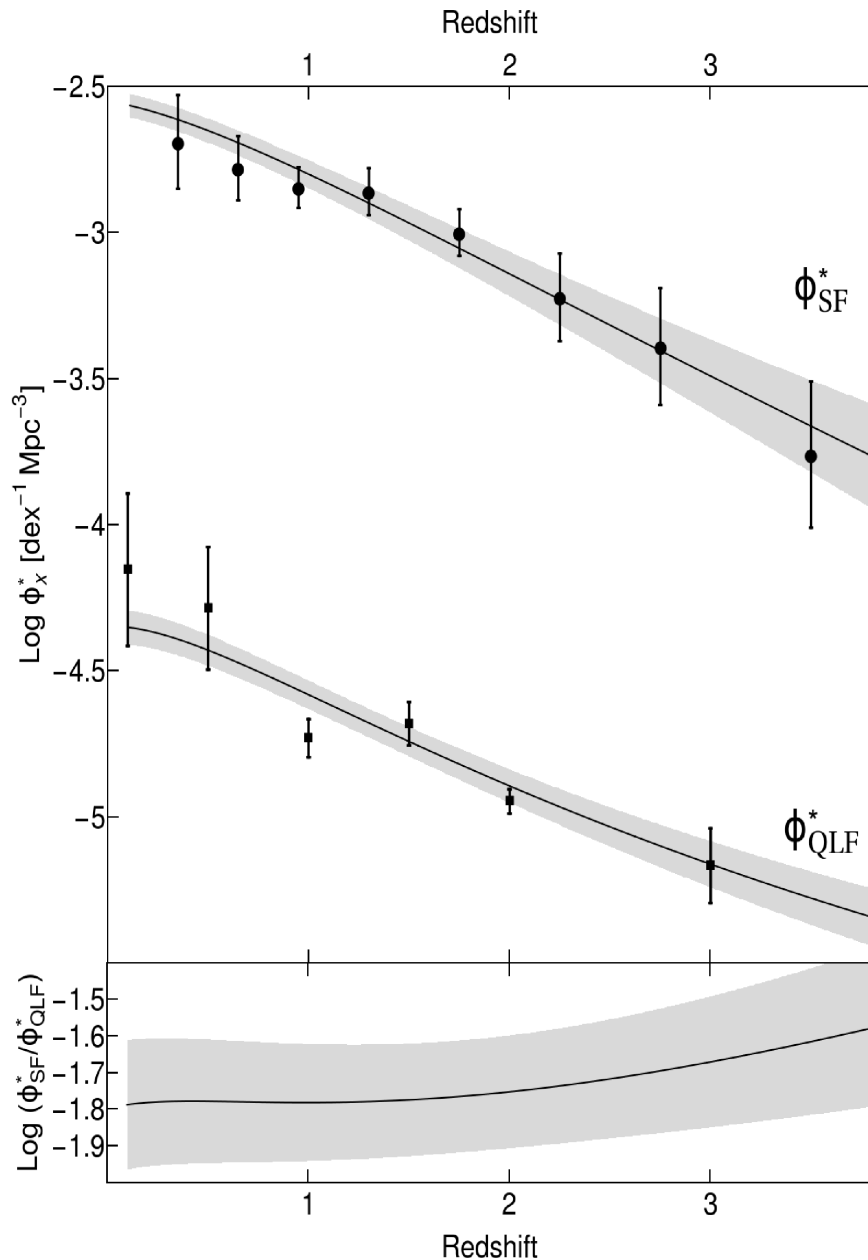
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- Normalization of quasar luminosity function changes as normalization of star forming galaxies

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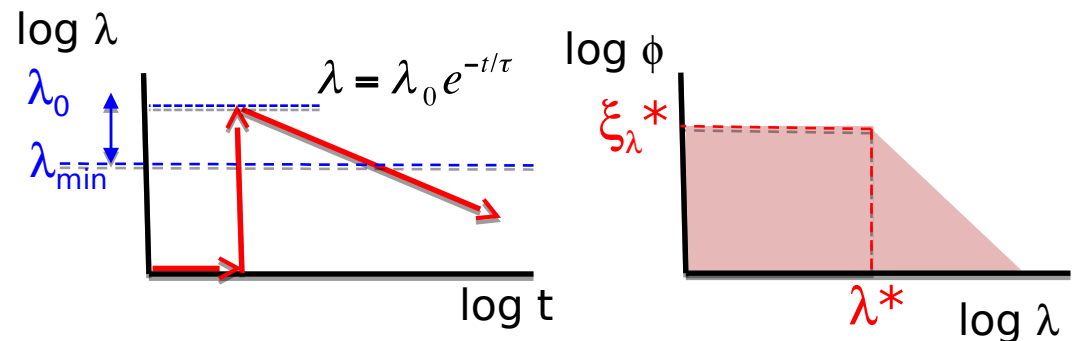
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- For example: in a “kick+decay” toy model
  - chance per unit time of kicking on,  $\eta$ ,
  - a distribution of size of kicks at  $\lambda_0 > \lambda_{\min}$
  - exponential decay time constant  $\tau$

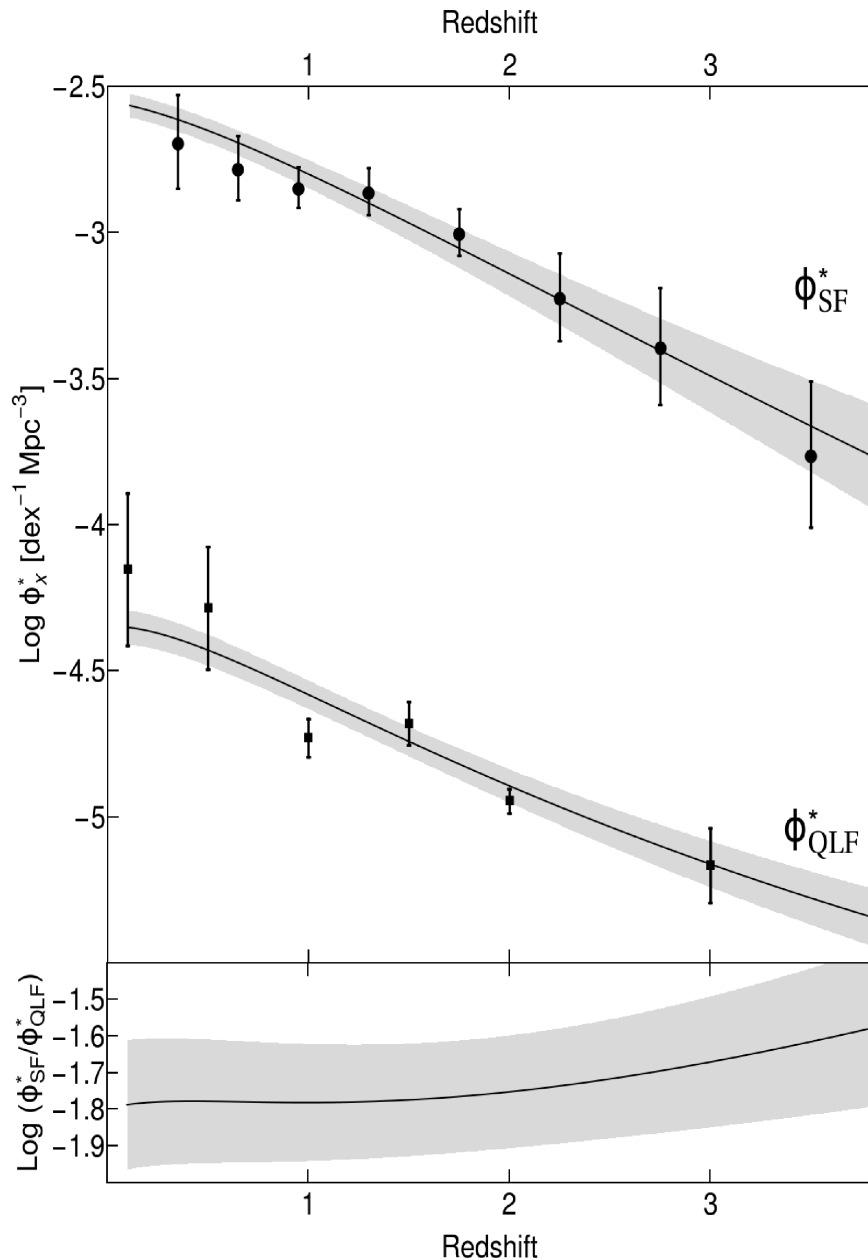
$$\xi_{\lambda}^* = \eta \tau$$

- Constant “duty cycle”





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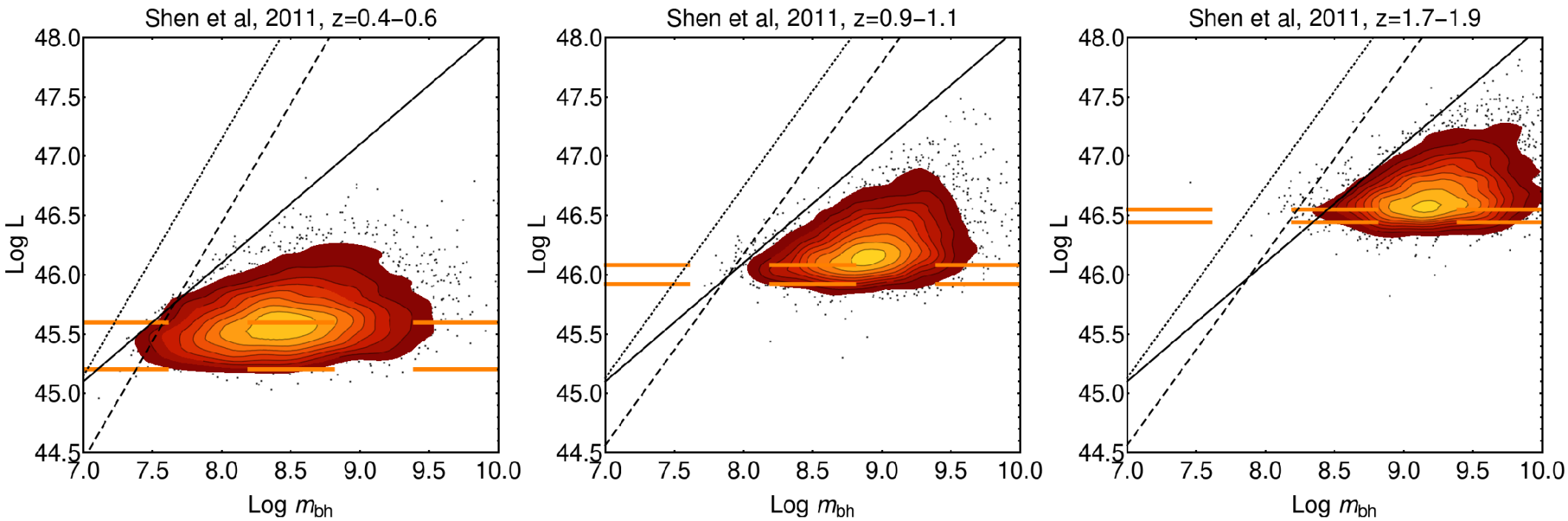
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- **Constant “duty cycle”**

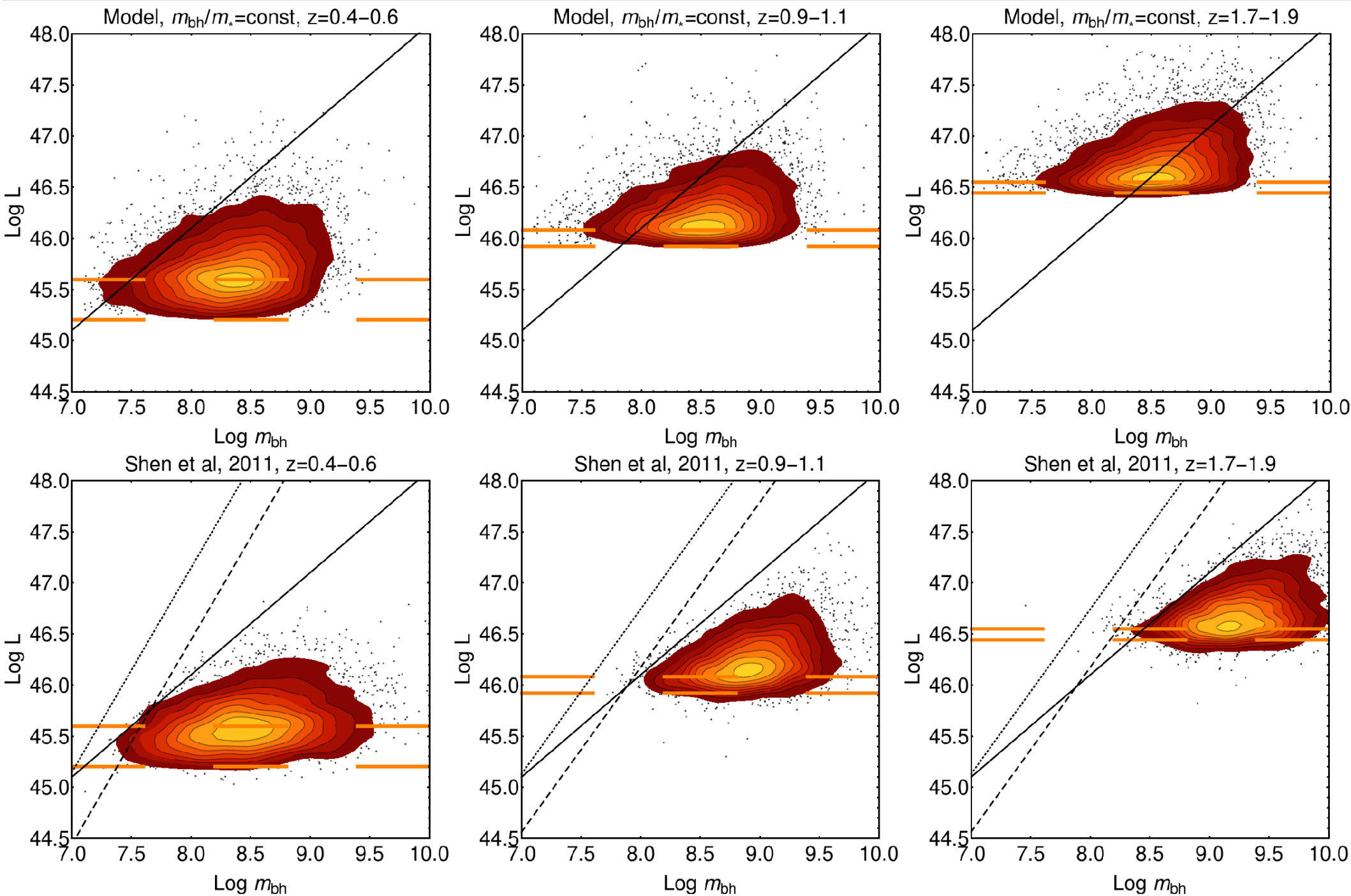
$$L^* \propto (1+z)^4 \quad z < 2$$

$$L^* \propto M^* m_{bh} / m_* \lambda^*$$

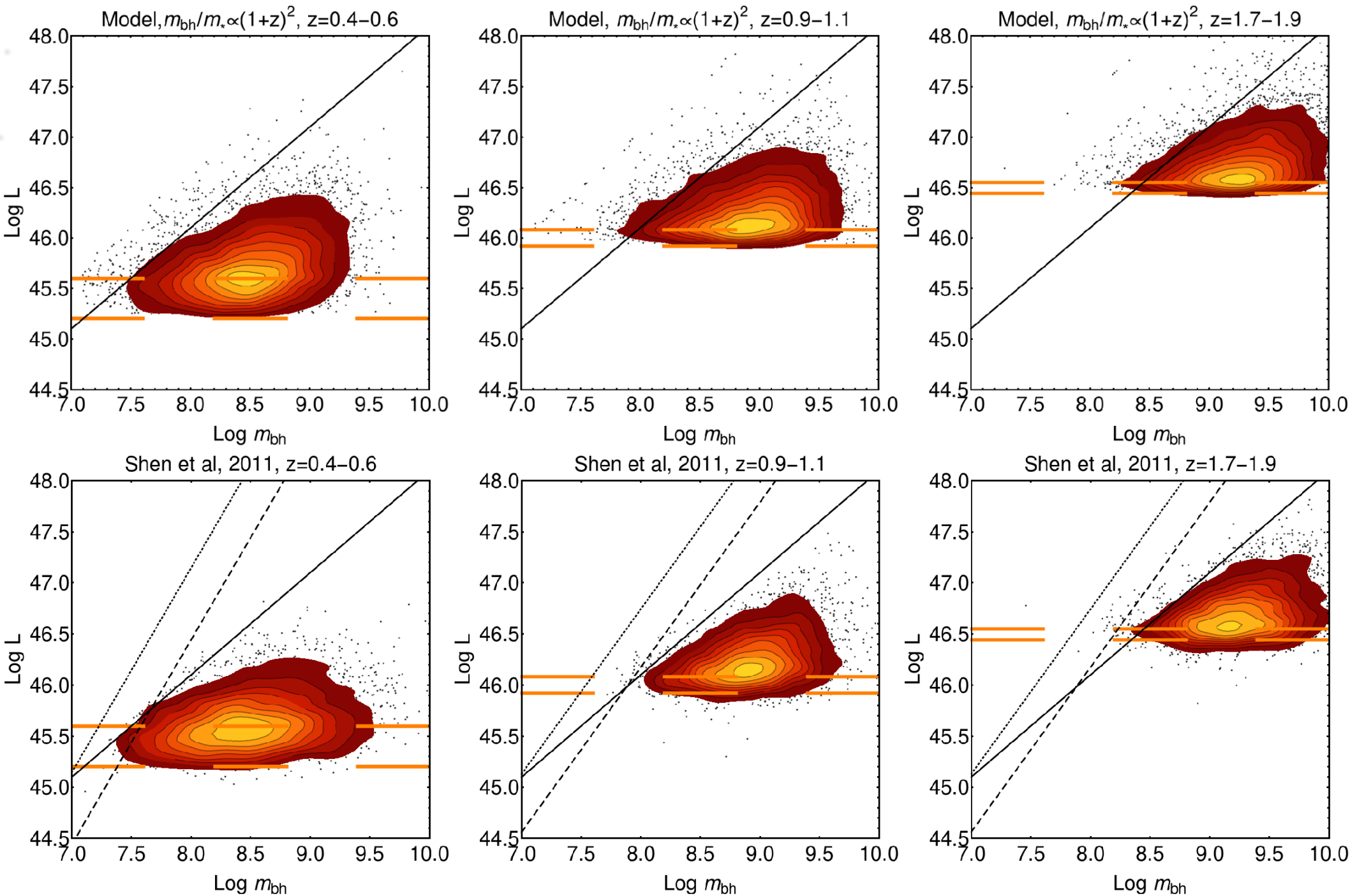
# Results from simulating mass-luminosity plane



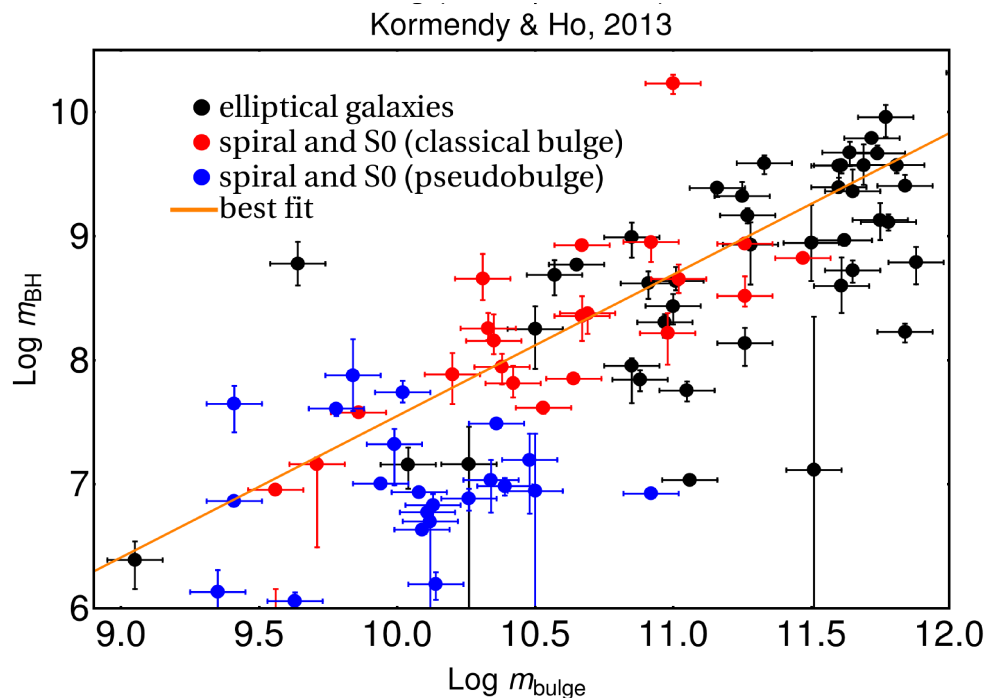
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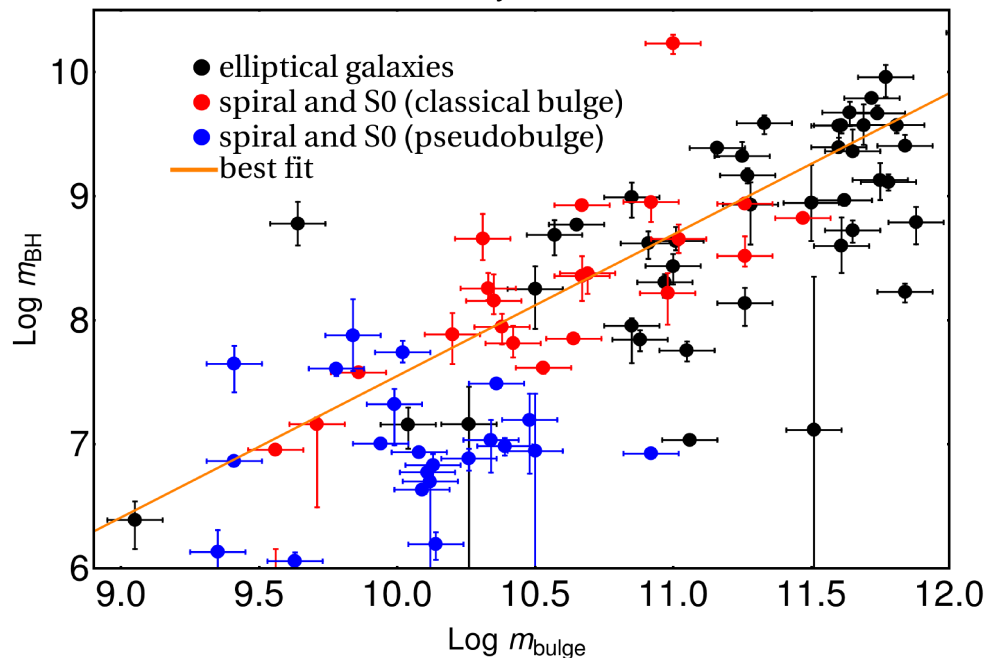
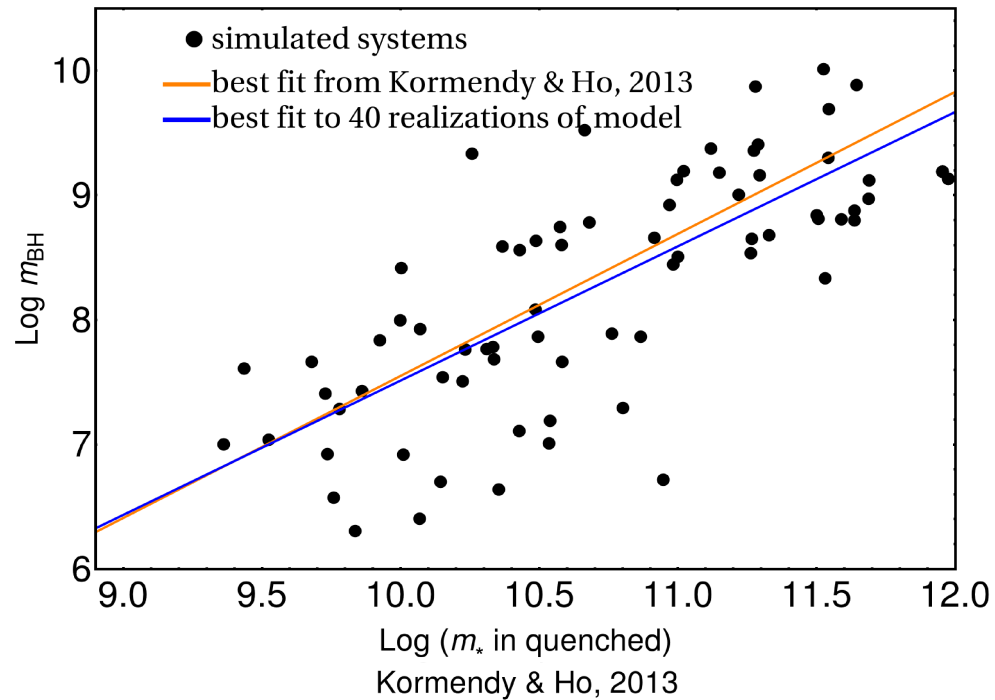


# Local relation is reproduced?



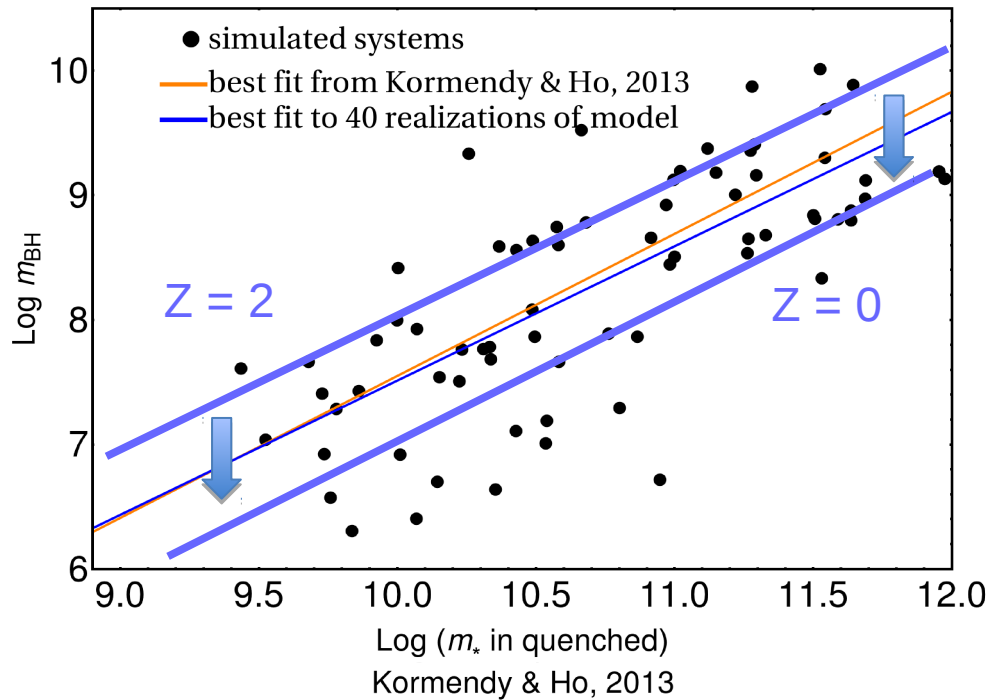
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model, quenching after  $z=2$

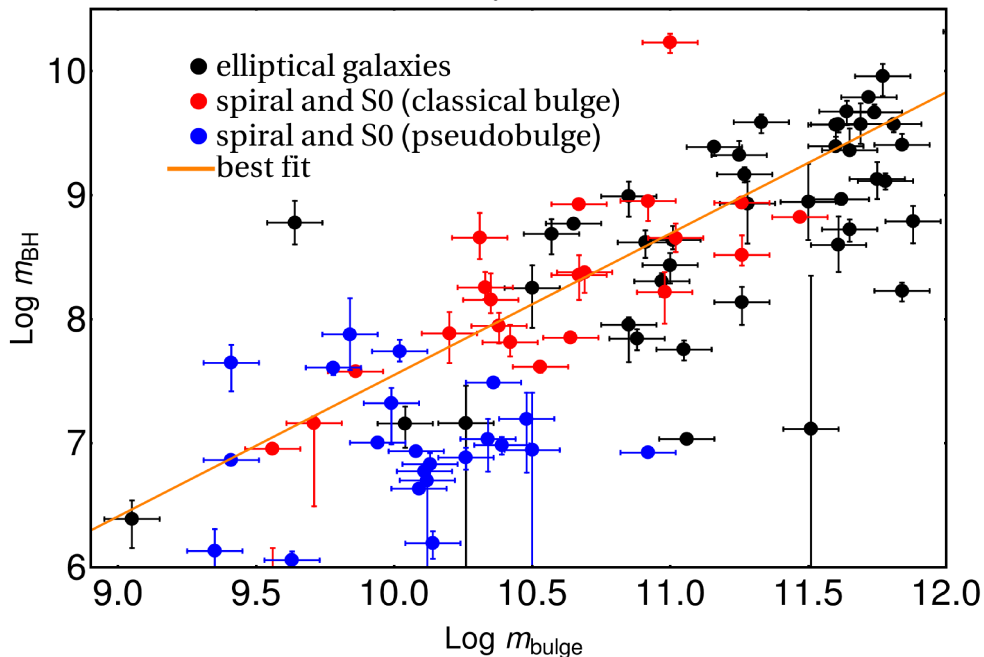


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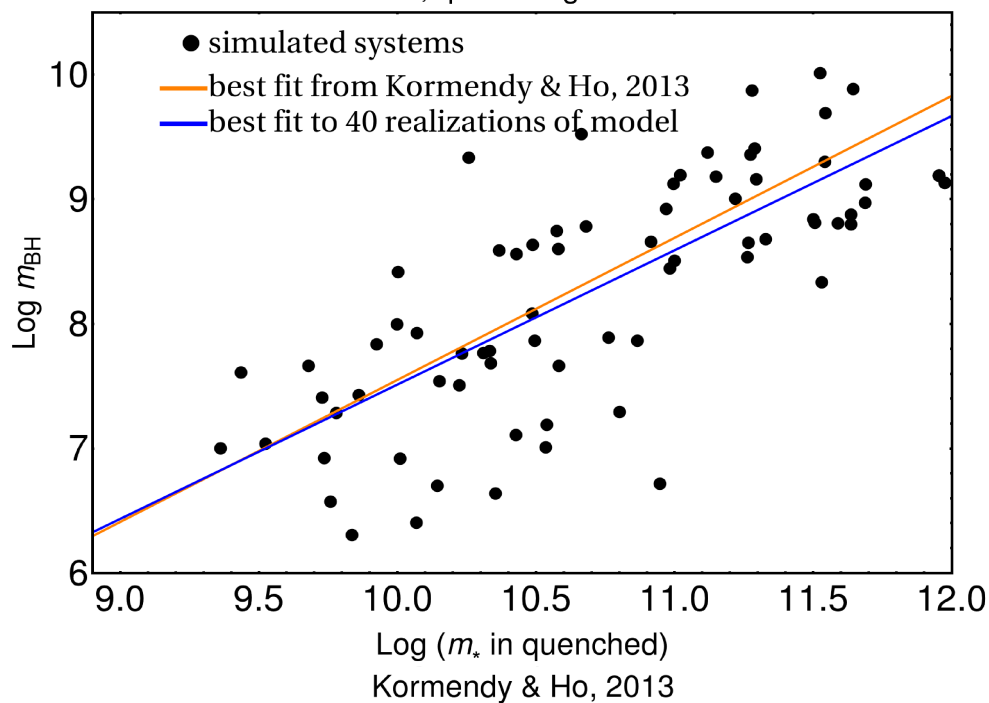


- Mean redshift of quenching for quenched galaxies we see today is at around redshift of 1 to 1.5.
- Galaxies which have quenched at low redshift will be below relation (pseudobulges?)
- Tilt in the relation

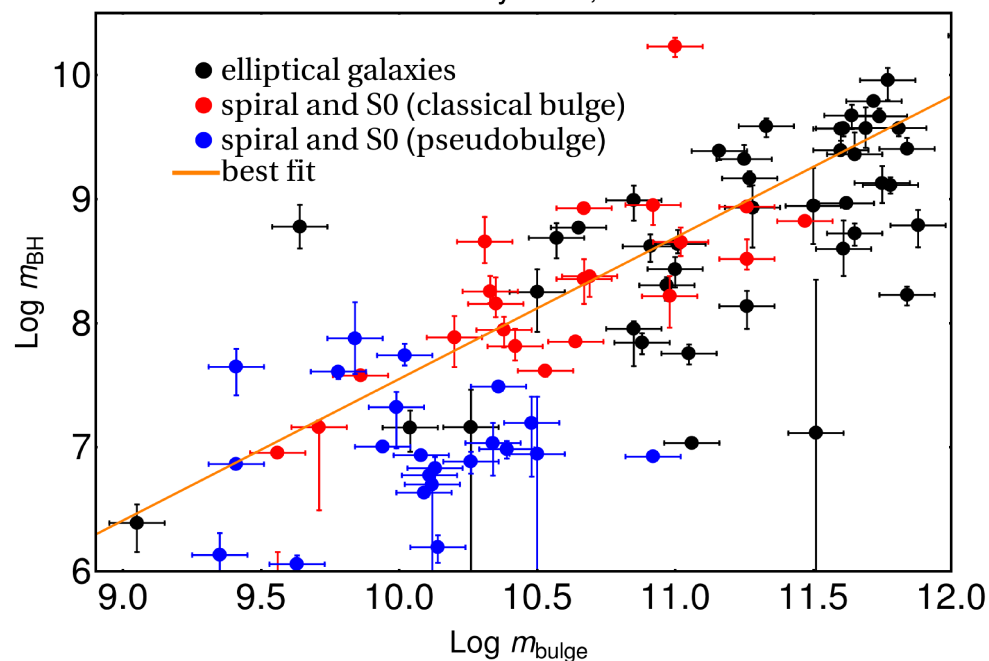


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model, quenching after z=2



Kormendy & Ho, 2013



- At a given stellar mass, the size of star-forming galaxies scales roughly as  $(1+z)^{-1}$
- $m_{bh} \propto 3 \cdot 10^8 \sigma_{200}^4$
- At a given galaxy mass  $r \propto (1+z)^{-1} \Leftrightarrow \sigma^2 \propto (1+z)$

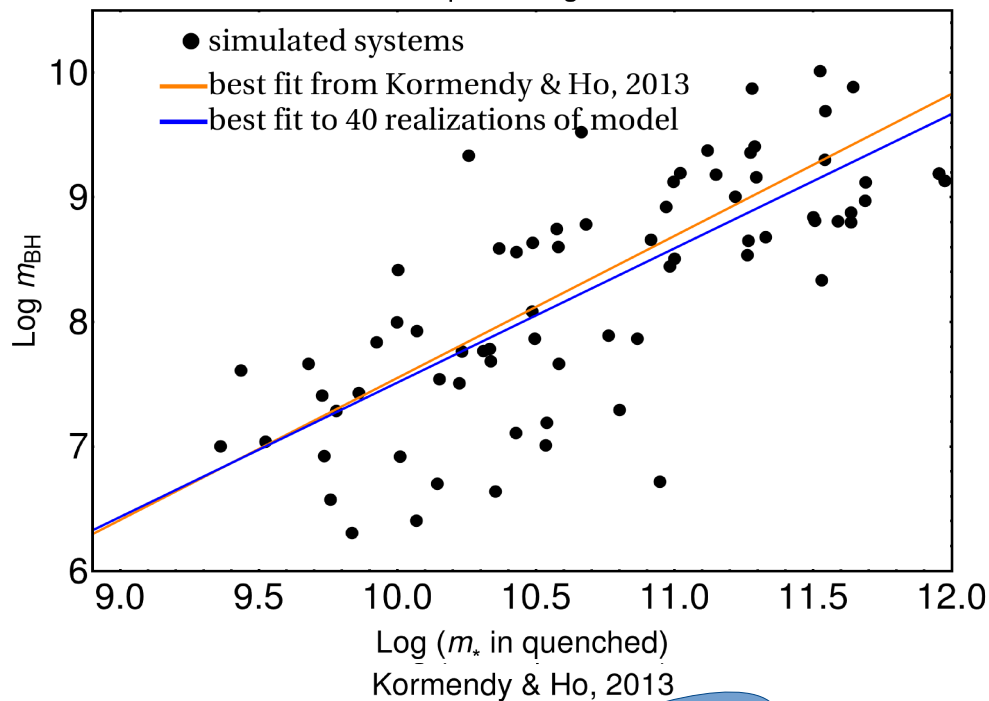
$$\frac{m_{bh}}{m_{star}} \propto (1+z)^2 \Leftrightarrow \frac{m_{bh}}{\sigma^4} = \text{constant}$$

- Constant  $m_{bh}$  – sigma, virial relation, and size evolution of galaxies lead to evolution in  $m_{bh}/m_*$

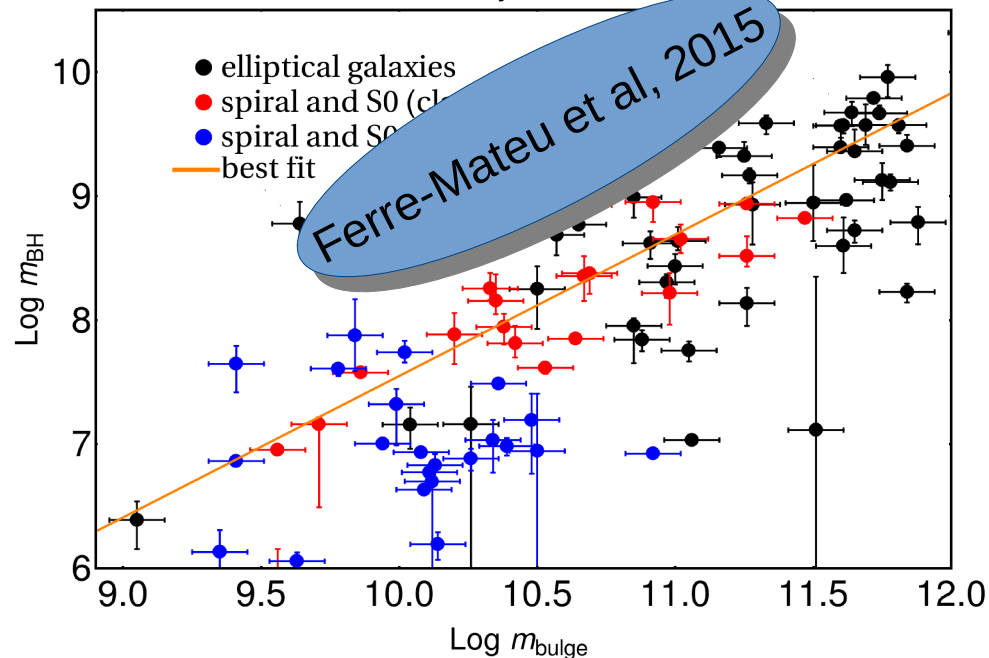


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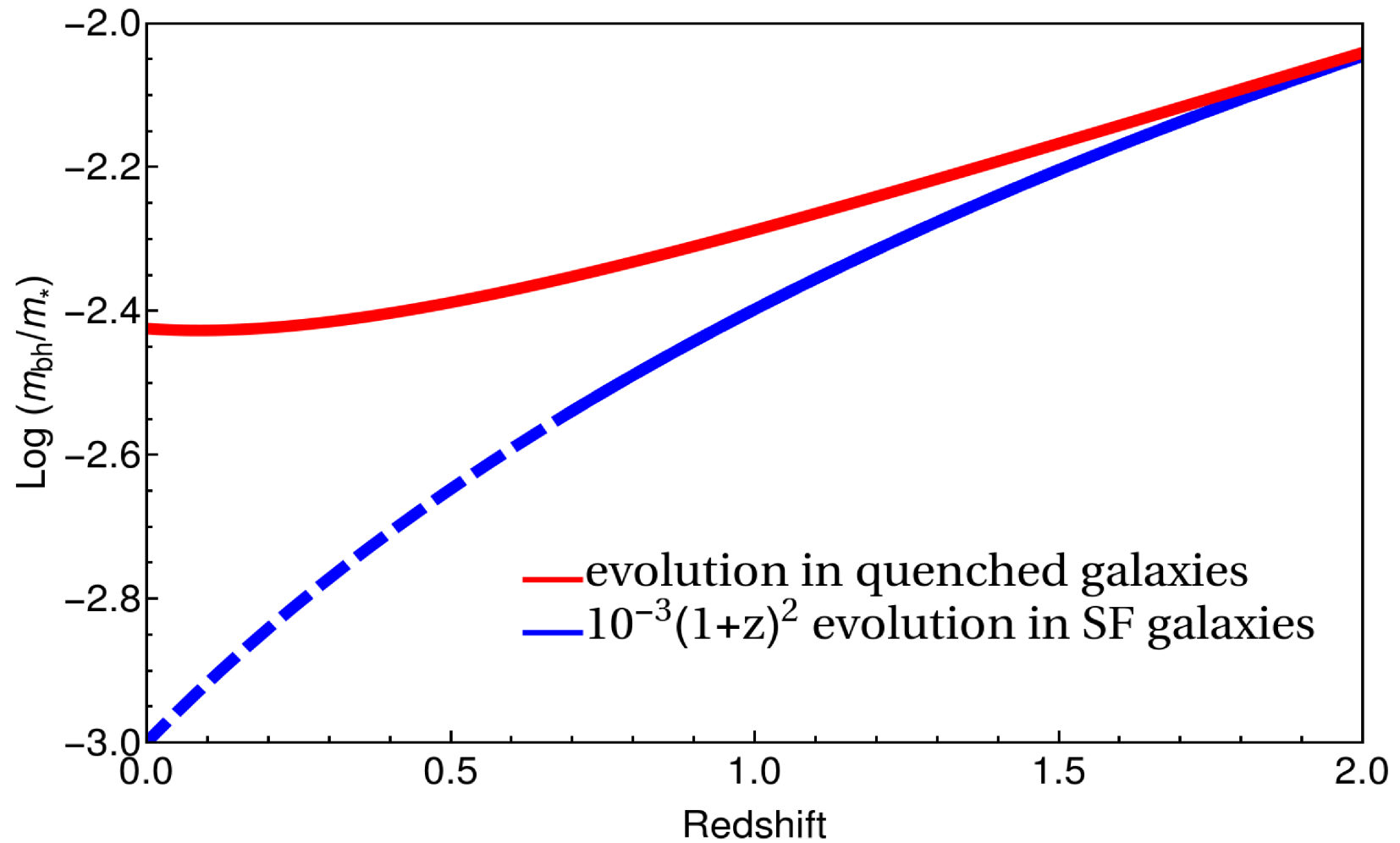


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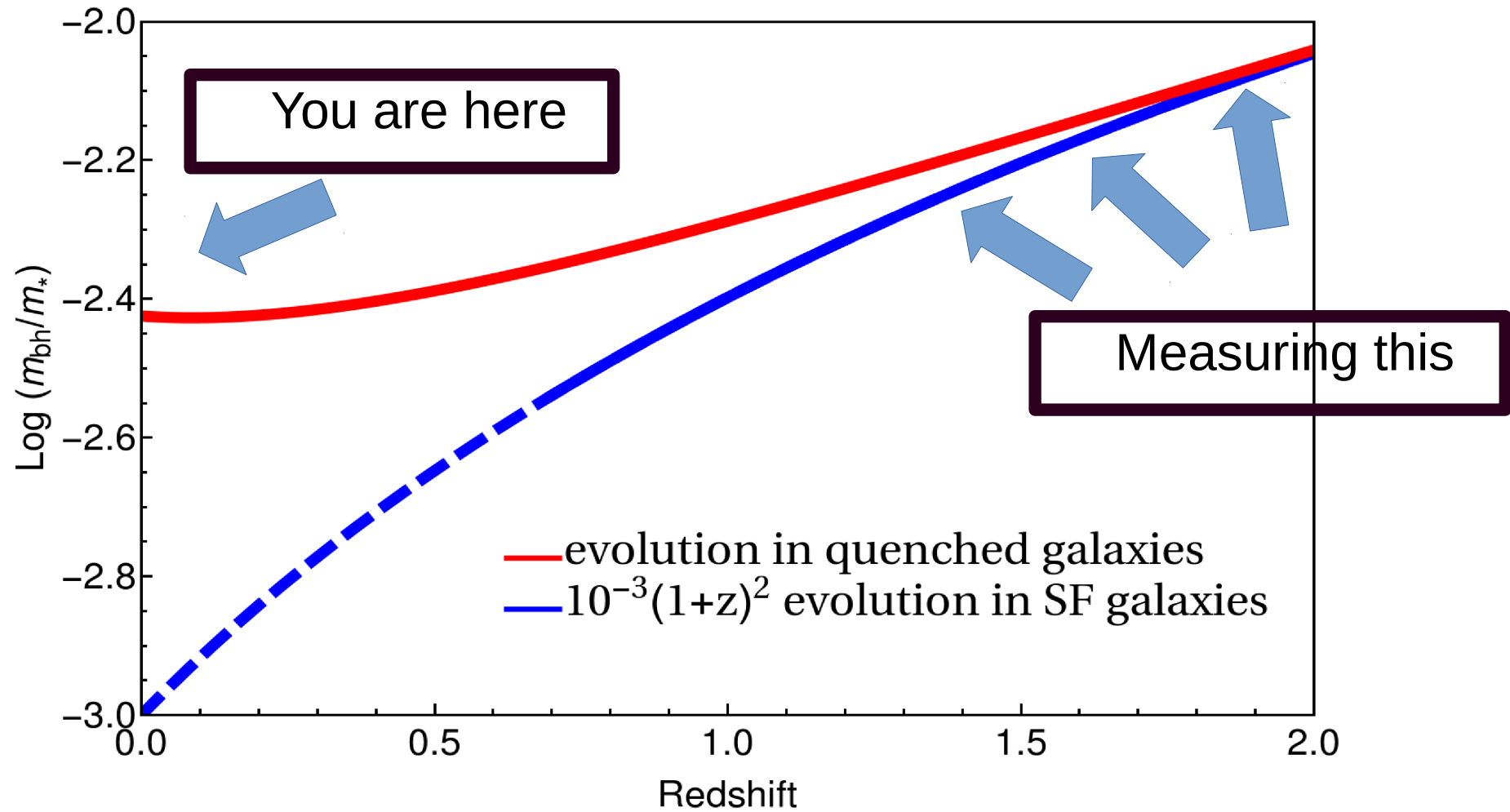
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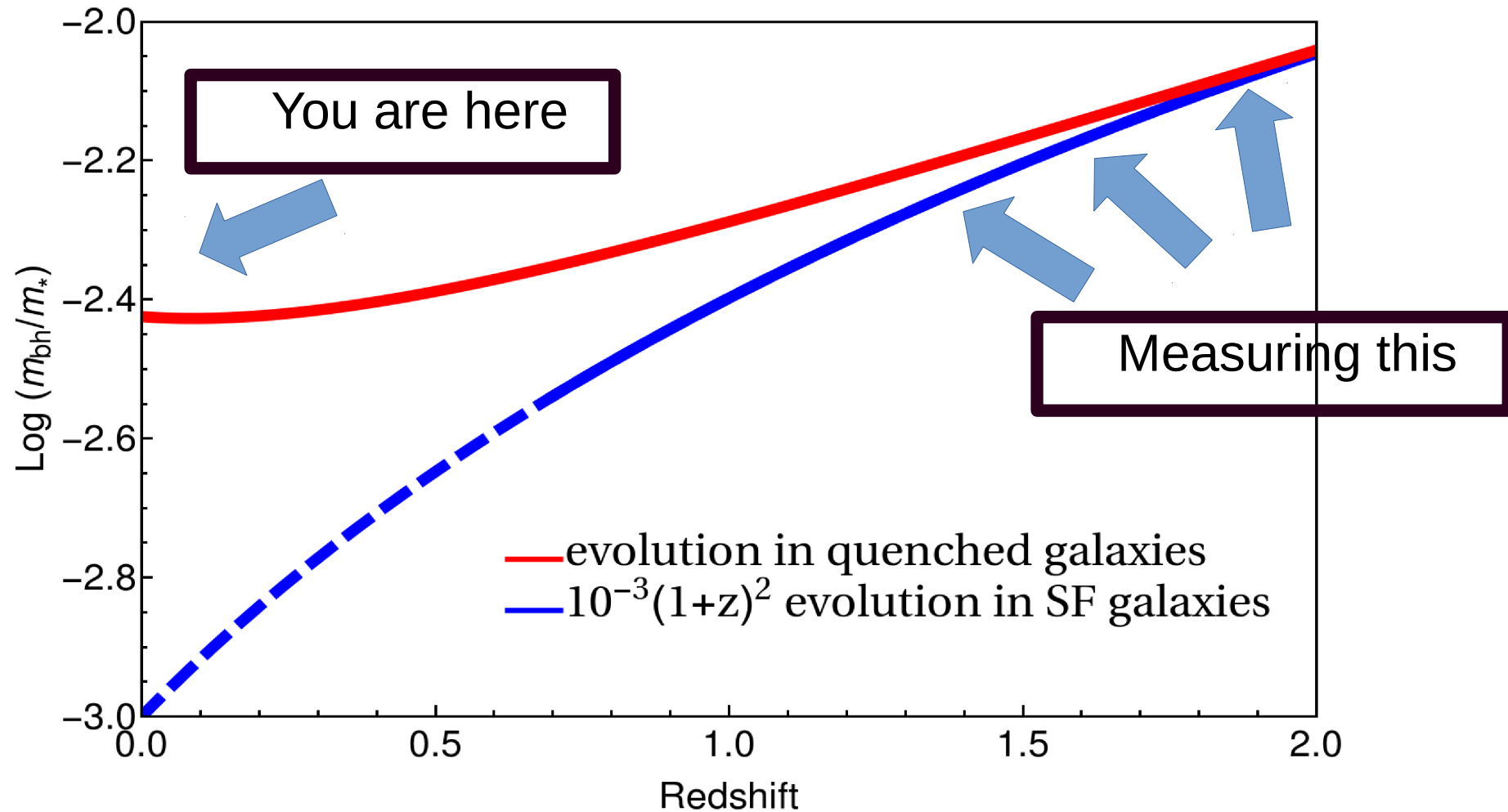
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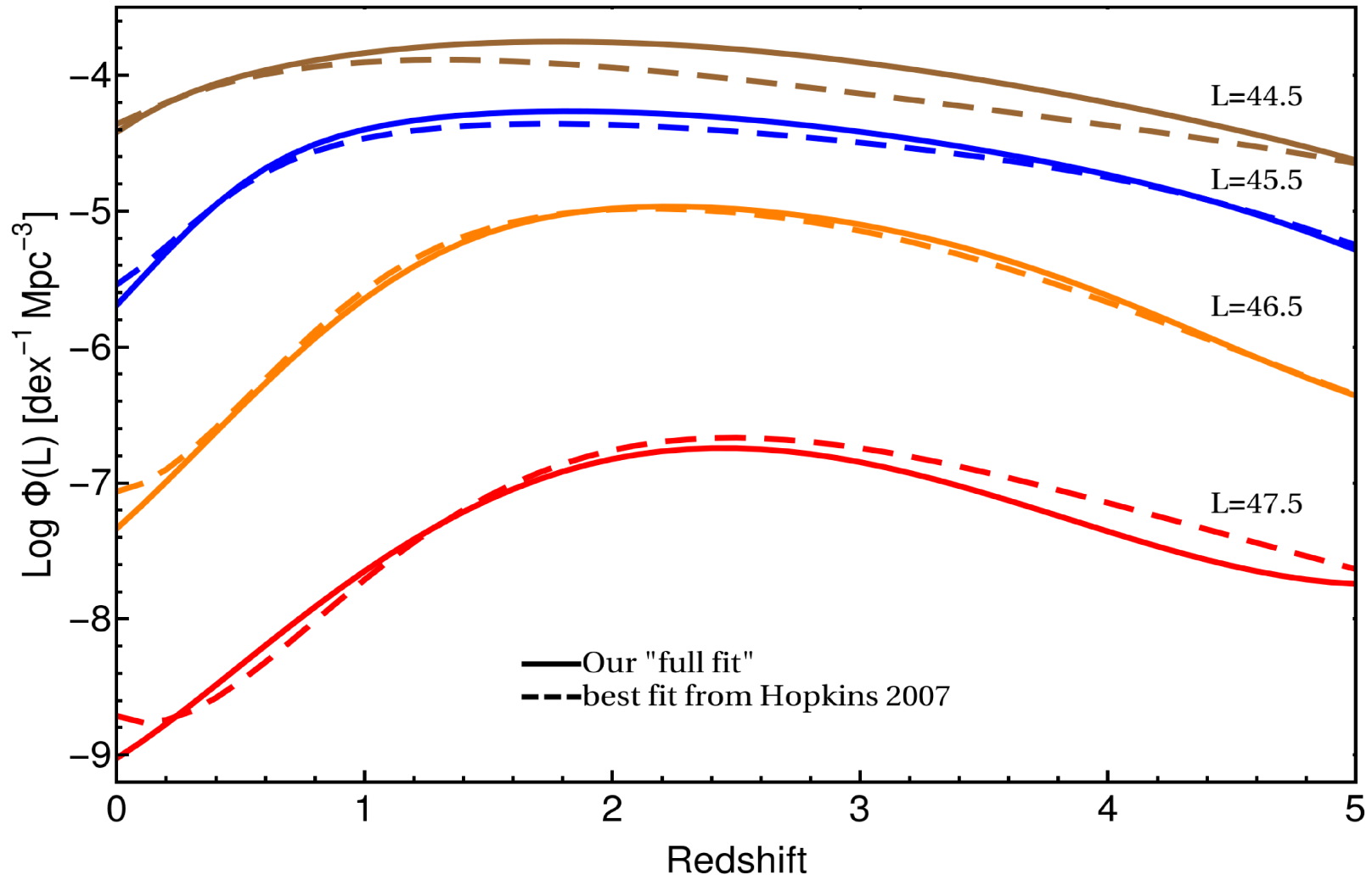


# Bias in measuring $m_{\text{bh}}/m_*$



- Measuring mass ratio in star-forming systems and comparing to local relation is potentially very dangerous

# What/where is downsizing



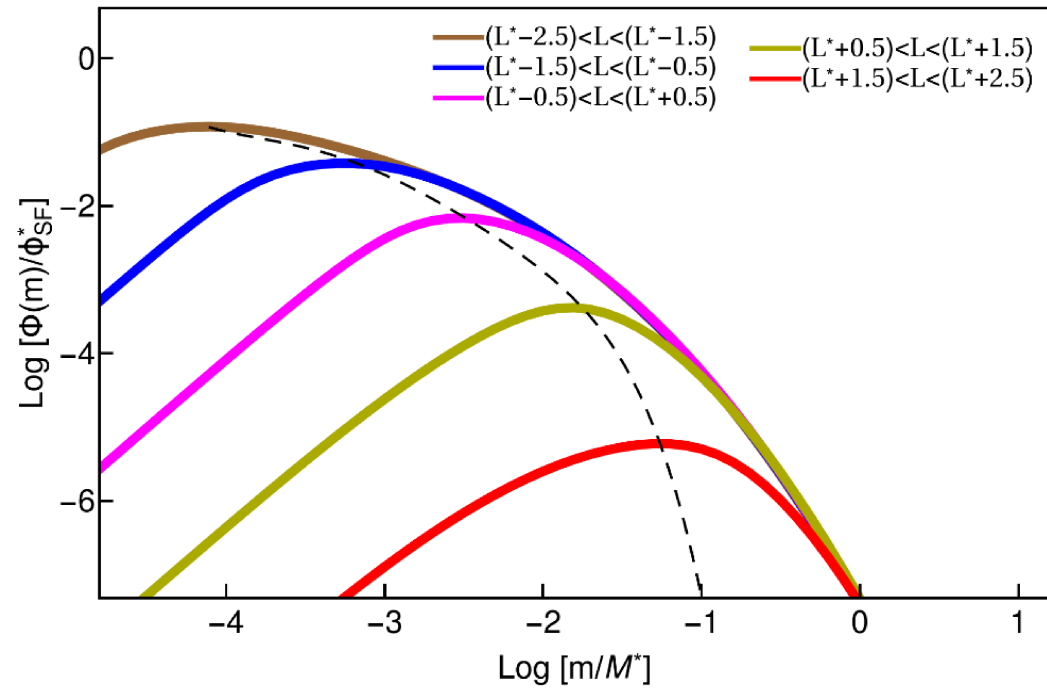
- “Downsizing” is reproduced even though Eddington ratio distribution is strictly mass-independent

# Summary

- Simple global model combining galaxy mass function and quasar luminosity function leads to following conclusions
  - **Constant “duty cycle”** at characteristic Eddington ratio
    - Evolution of  $\phi_{QLF}^*$  normalization of quasar luminosity function is consistent with  $\phi_{SF}^*$  normalization of star-forming galaxies
  - **Evolution in the  $m_{bh}/m_*$  relation** in star-forming galaxies
    - Evolution in  $L^*$  can be due to evolution of  $\lambda^*$  and/or  $m_{bh}/m_*$
    - Non-evolving  $m_{bh}/m_*$  disfavored by mass-luminosity plane
    - Local relation and measurements at higher redshift are satisfied by evolving relation
    - Size evolution in galaxies implies evolution in either  $m_{bh}/m_*$  or  $m_{bh} - \sigma$  relation
    - Extreme caution when comparing black holes in star-forming and quenched galaxies

# Additional slides

Black hole mass function



Black hole host galaxy mass function

