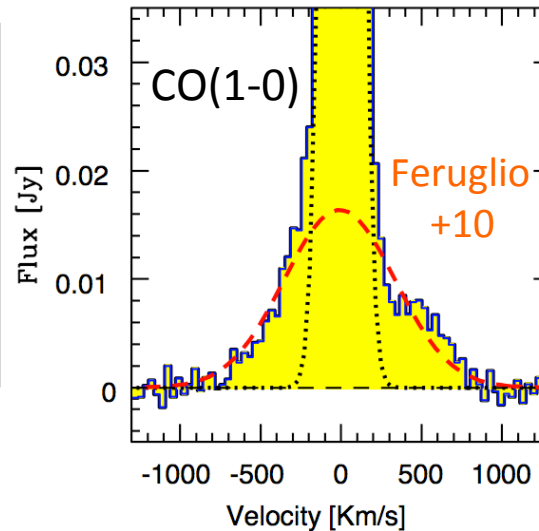
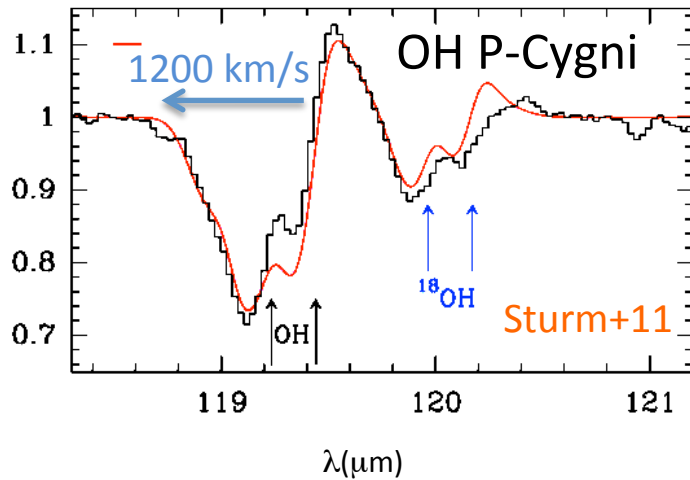


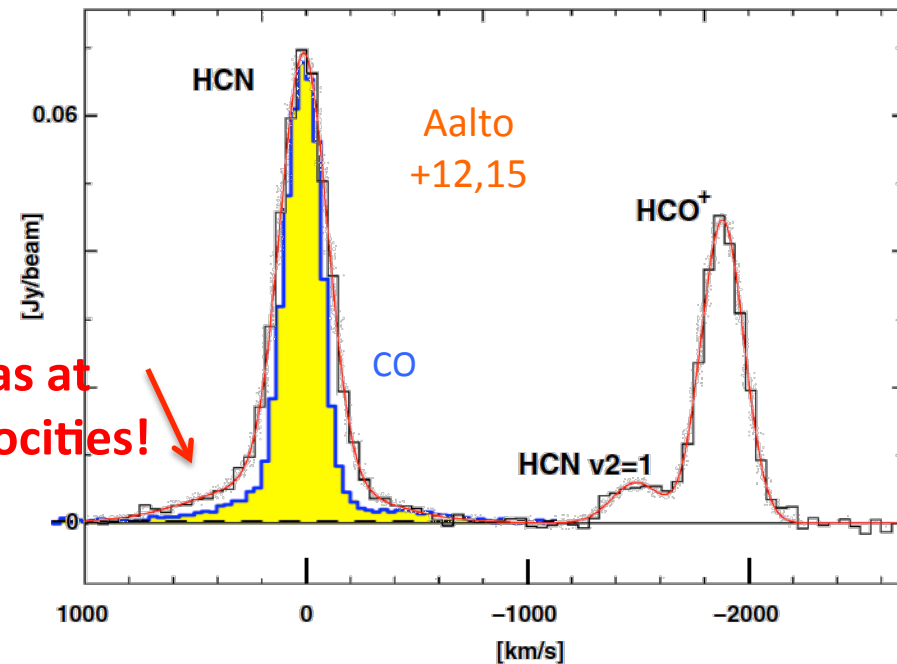
Powerful AGNs (QSOs) can seriously “hurt” their host galaxies

massive AGN-driven molecular outflows

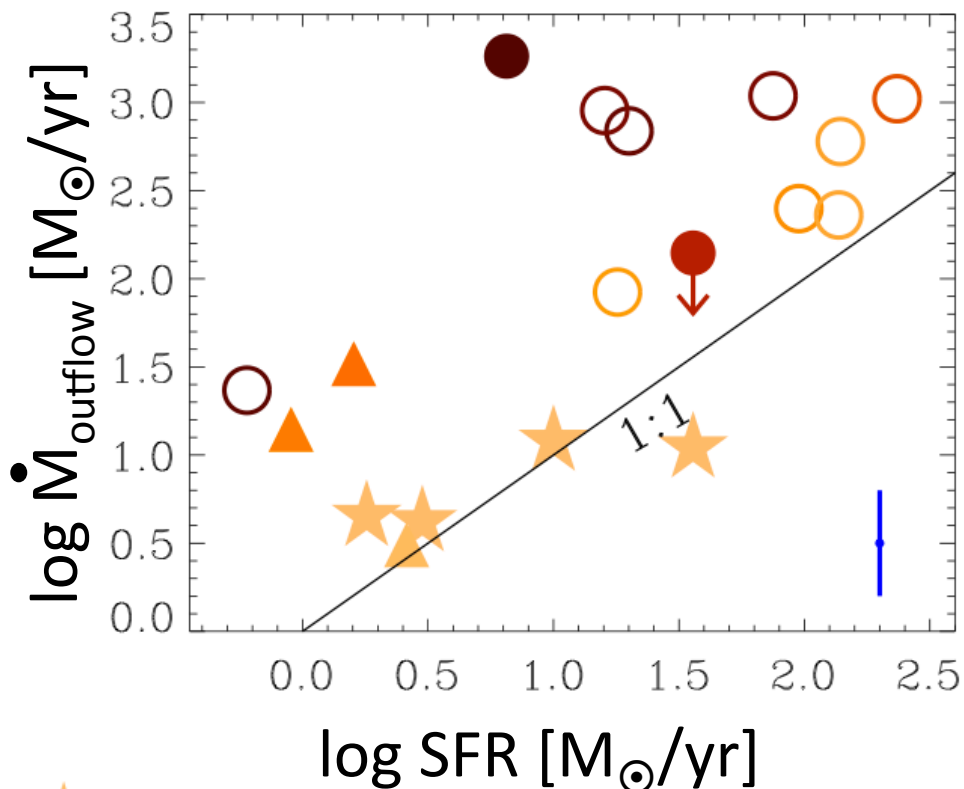


even outflows of large amount of DENSE molecular gas

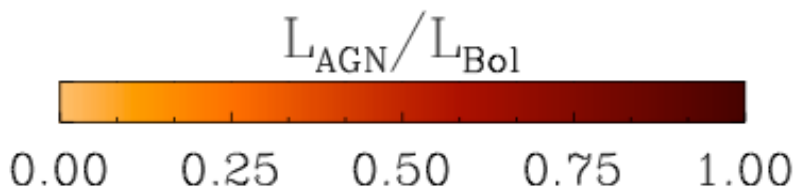
dense gas at high velocities!



Outflow rate can be boosted by a large factor by an AGN

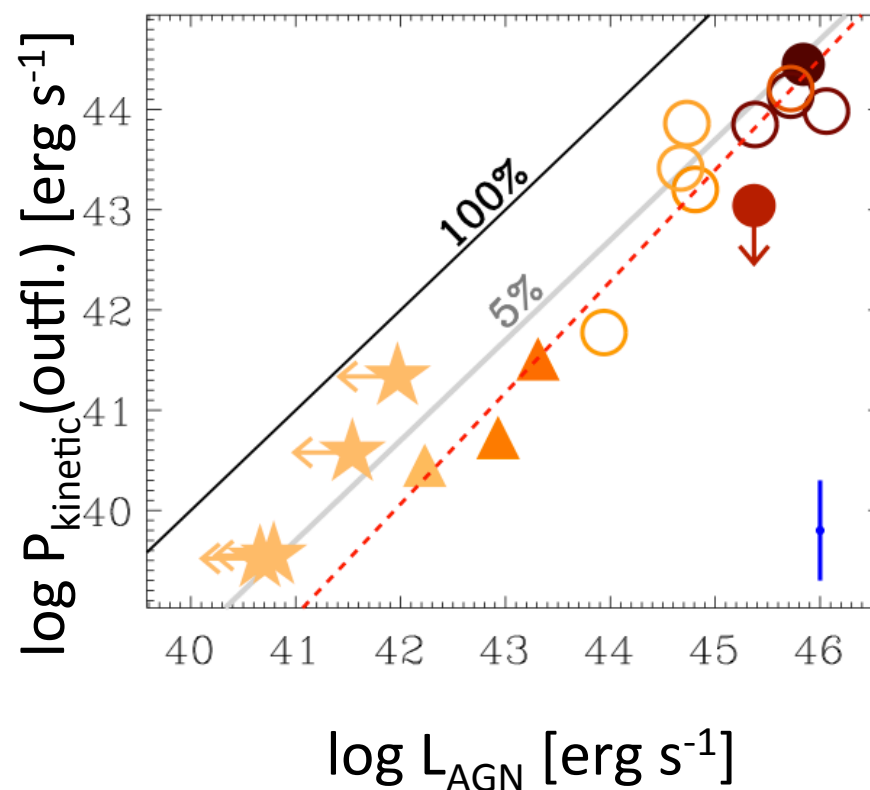


- ★ SB
- Sy2
- Sy1
- ▲ LINER



Cicone+13

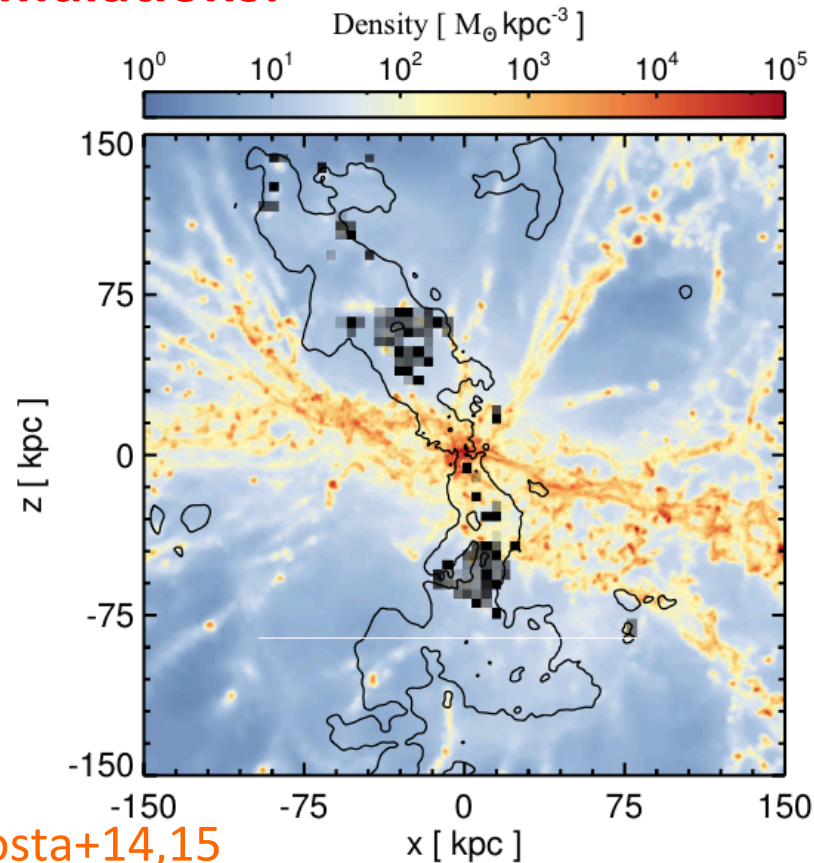
Kinetic power & momentum rate in agreement with AGN-driven winds models



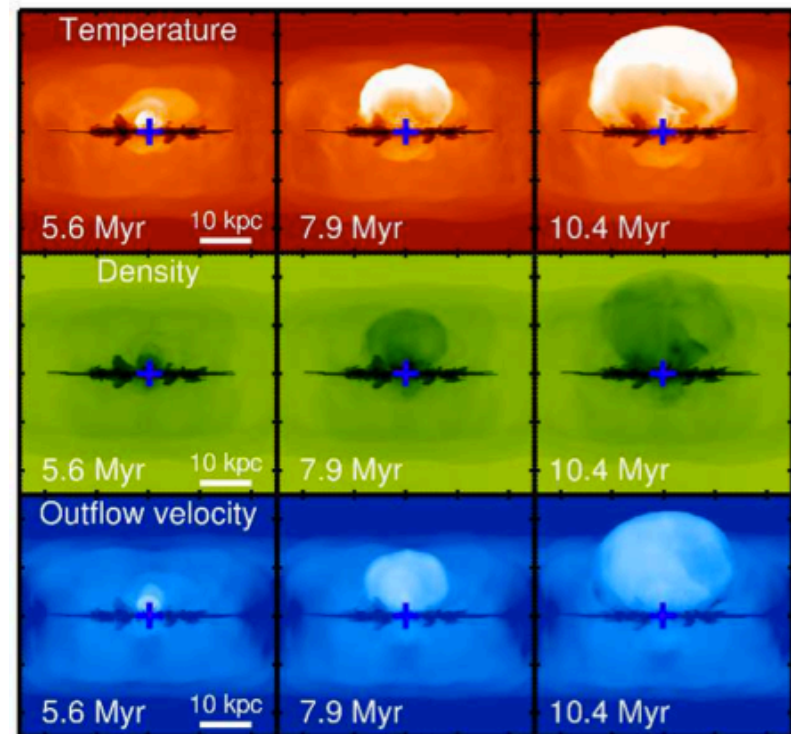
In terms of energy & momentum rate AGNs can quench the host galaxy and stop BH accretion
-> BH-sigma relation [King+10, etc.]

Can AGN-driven outflows really quench star formation over the entire galaxy, or even in its central region ?

Simulations:



Costa+14,15



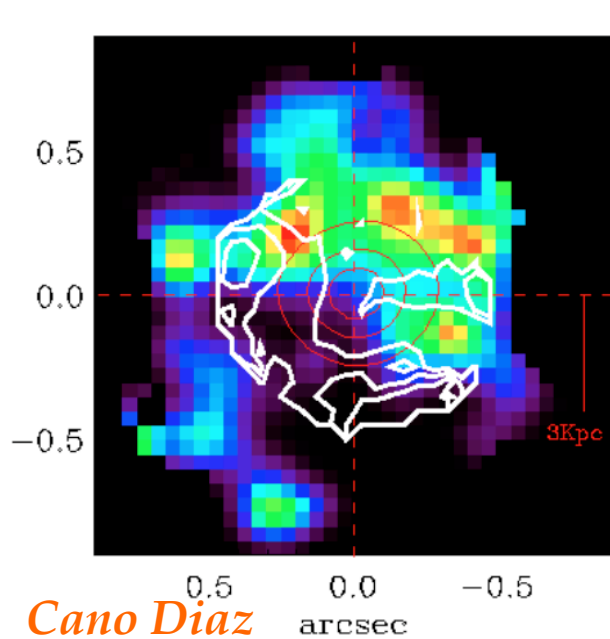
Gabor+14, Roos+15

- Outflowing gas mostly escapes through low density, least resistance regions
- Does not stop inflow
- Does not escape the halo, rains back onto the galaxy

Observations:

Outflows in the most powerful quasars at $z \sim 1-2$

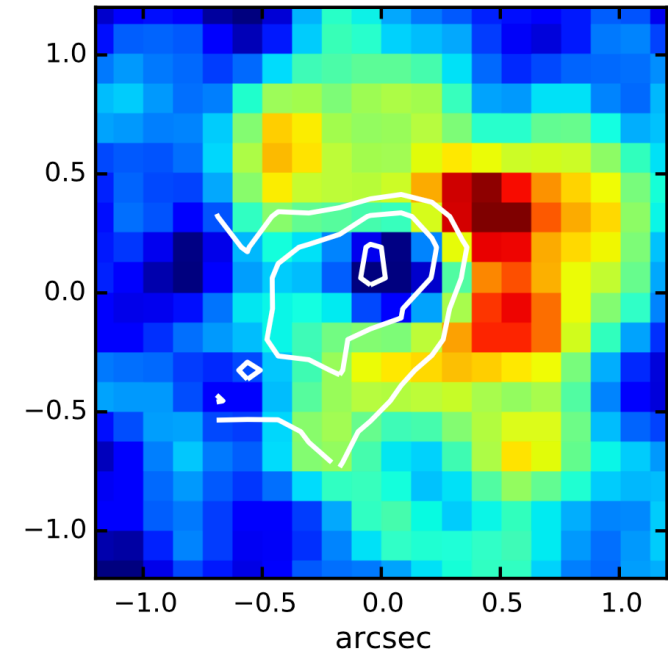
(epoch when quenching in most massive galaxies must occur)



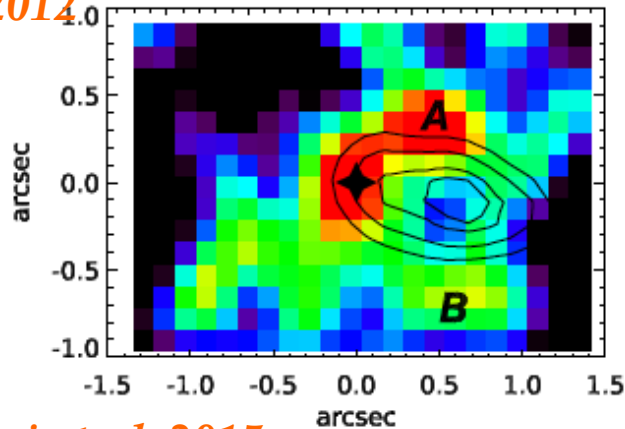
Color: narrow H α = SFR

Contours: high velocity
gas (outflow) from [OIII]

*Cano Diaz
et al. 2012*



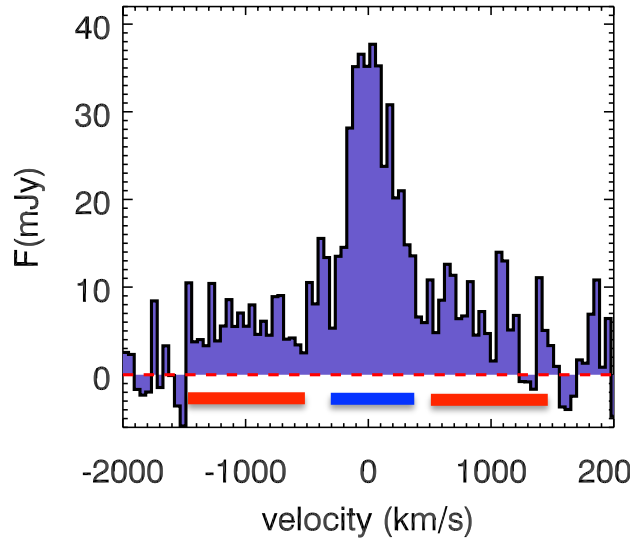
Carniani et al. 2015



Cresci et al. 2015

Star formation is suppressed in those regions affected by the quasar driven outflow, but star formation proceeds happily (at $\sim 100 M_{\text{sun}}/\text{yr}$) elsewhere

Outflows in the most luminous QSOs at $z \sim 6$ traced with [CII]158 μm

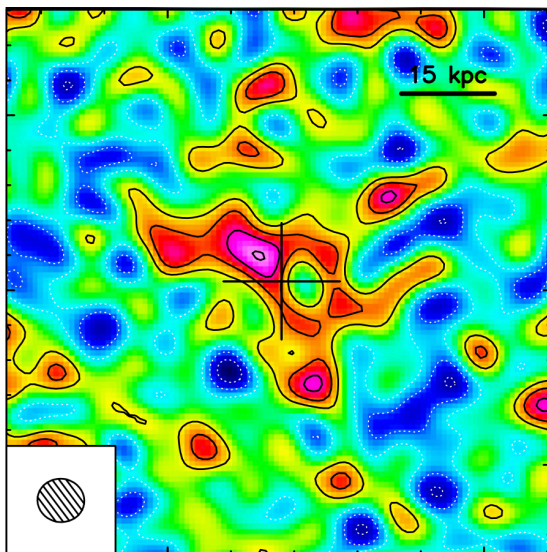
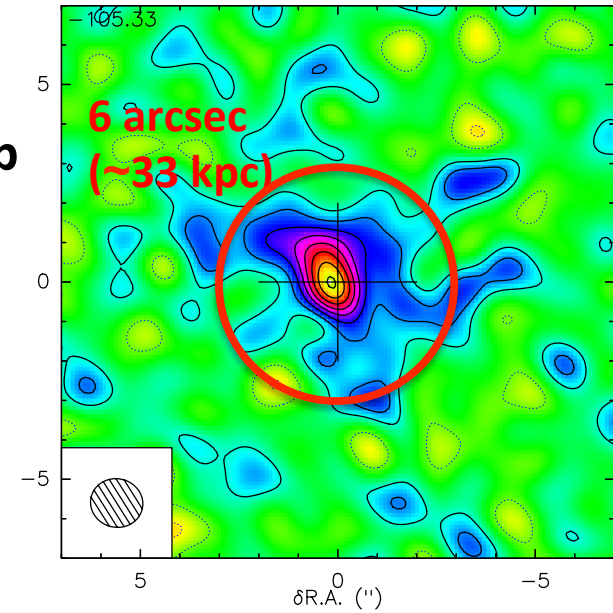


Velocity
integrated map



Maiolino+12

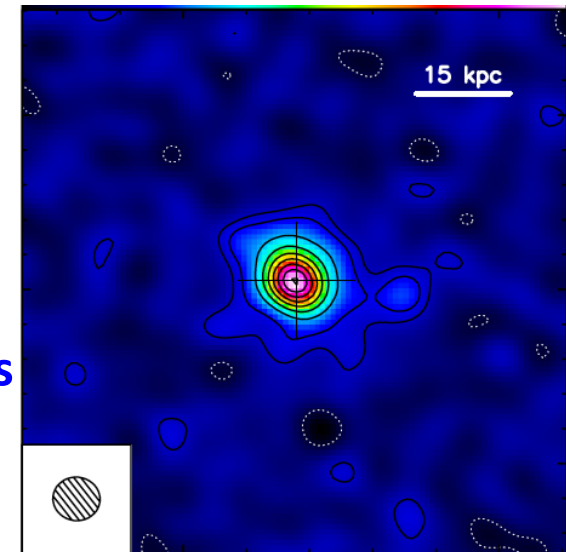
Cicone+15



High velocity gas: outflow
extending out to 30 kpc!

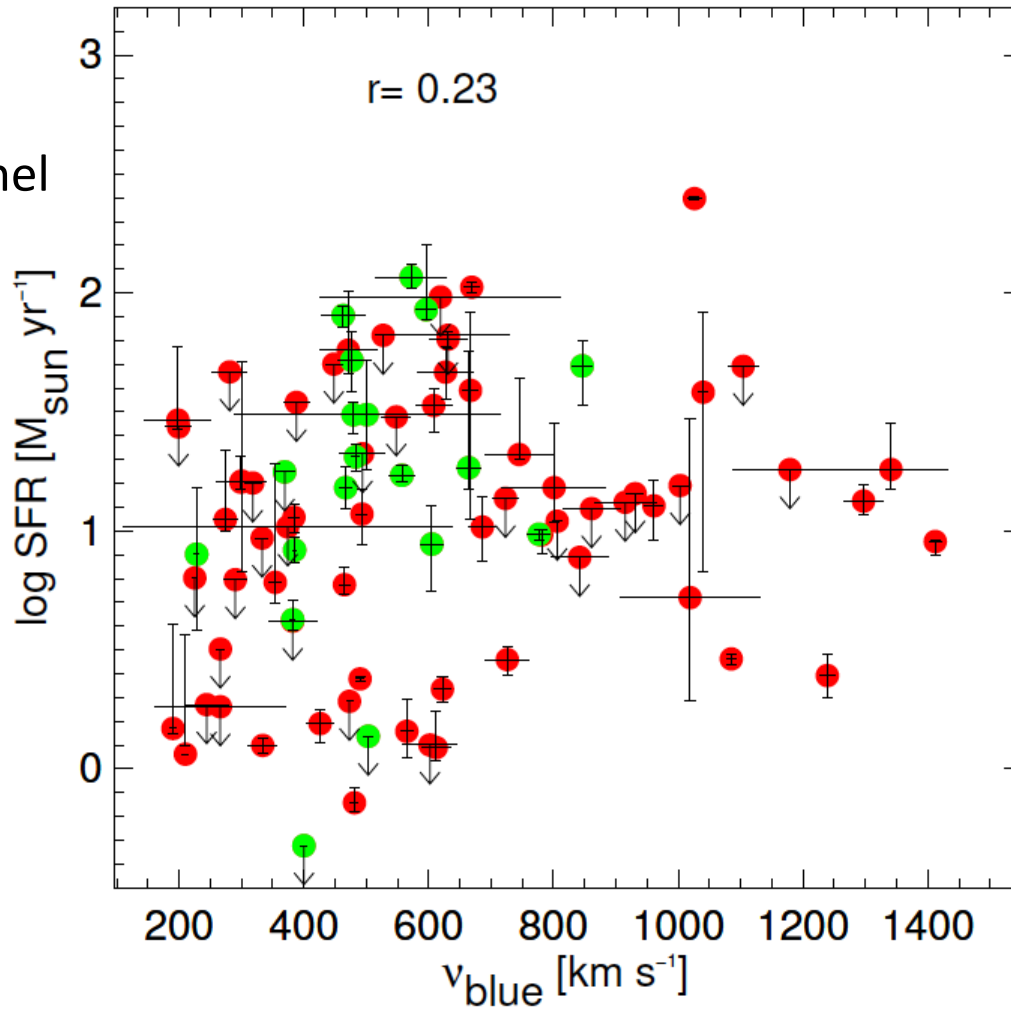


Low velocity gas: mostly
within 1.5kpc forming stars
at 1000 M_{sun}/yr



Low redshift quasars ($z < 1$)

SFR from
FIR-Herschel



No correlation between
outflow velocity and SFR

Outflow velocity from [OIII] profile

Balmaverde+15